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# A Note on LP-Kenmotsu Manifolds Admitting Conformal Ricci-Yamabe Solitons

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Abstract. In the current note, we study Lorentzian para-Kenmotsu (in brief, *LP*-Kenmotsu) manifolds admitting conformal Ricci-Yamabe solitons (CRYS) and gradient conformal Ricci-Yamabe soliton (gradient CRYS). At last by constructing a 5-dimensional non-trivial example we illustrate our result.

### 1. Introduction

As a generalization of the classical Ricci flow [8], the concept of conformal Ricci flow was introduced by Fischer [5], which is defined on an n-dimensional Riemannian manifold M by the equations

$$\frac{\partial g}{\partial t} = -2(S + \frac{g}{n}) - pg, \quad r(g) = -1,$$

where p defines a time dependent non-dynamical scalar field (also called the conformal pressure), g is the Riemannian metric; r and S represent the scalar curvature and the Ricci tensor of M, respectively. The term -pg plays a role of constraint force to maintain r in the above equation.

In [1], the authors Basu and Bhattacharyya proposed the concept of conformal Ricci soliton on M and is defined by

$$\pounds_{\mathcal{K}}g+2S+(2\Lambda-(p+\frac{2}{n}))g=0,$$

where  $\mathcal{L}_{\mathcal{K}}$  represents the Lie derivative operator along the smooth vector field  $\mathcal{K}$  on  $\mathcal{M}$  and  $\Lambda \in \mathbb{R}$  (the set of real numbers).

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Very recently, a scalar combination of Ricci and Yamabe flows was proposed by the authors Güler and Crasmareanu [7], this advanced class of geometric flows called Ricci-Yamabe (RY) flow of type  $(\sigma, \rho)$  and is defined by

$$\frac{\partial}{\partial t}g(t) + 2\sigma S(g(t)) + \rho r(t)g(t) = 0, \quad g(0) = g_0$$

for some scalars  $\sigma$  and  $\rho$ . A solution to the RY flow is called Ricci-Yamabe soliton (RYS) if it depends only on one parameter group of diffeomorphism and scaling. A Riemannian (or semi-Riemannian) manifold *M* is said to have a RYS if [9, 10]

$$\pounds_{\mathcal{K}}g + 2\sigma S + (2\Lambda - \rho r)g = 0. \tag{1.1}$$

A Riemannian (or semi-Riemannian) manifold M is said to have a conformal Ricci-Yamabe soliton (CRYS) if [20]

$$\pounds_{\kappa}g + 2\sigma S + (2\Lambda - \rho r - (p + \frac{2}{n}))g = 0,$$
 (1.2)

where  $\sigma$ ,  $\rho$ ,  $\Lambda \in \mathbb{R}$ .

If K is the gradient of a smooth function v on M, then (1.2) is called the gradient conformal Ricci-Yamabe soliton (gradient CRYS) and hence (1.2) turns to

$$\nabla^2 v + \sigma S + (\Lambda - \frac{\rho r}{2} - \frac{1}{2}(\rho + \frac{2}{n}))g = 0, \qquad (1.3)$$

where  $\nabla^2 v$  is the Hessian of v and is defined by  $Hessv = \nabla \nabla v$ .

A CRYS is said to be shrinking, steady or expanding if  $\Lambda < 0, = 0$  or > 0, respectively. A CRYS is said to be a

- Conformal Ricci soliton if  $\sigma = 1$ ,  $\rho = 0$ ,
- Conformal Yamabe soliton if  $\sigma = 0, \rho = 1$ ,
- Conformal Einstein soliton if  $\sigma = 1, \rho = -1$ .

As a continuation of this study, we tried to study CRYS and gradient CRYS in the frame-work of *LP*-Kenmotsu manifolds of dimension *n*. We recommend the papers [2-4,6,13-17] and the references therein for more details about the related studies.

#### 2. Preliminaries

An *n*-dimensional differentiable manifold *M* with structure  $(\varphi, \zeta, \nu, g)$  is said to be a Lorentzian almost paracontact metric manifold, if it admits a (1, 1)-tensor field  $\varphi$ , a contravariant vector field  $\zeta$ , a 1-form  $\nu$  and a Lorentzian metric *g* satisfying

$$\nu(\zeta) + 1 = 0, \tag{2.1}$$

$$\varphi^2 E = E + \nu(E)\zeta, \qquad (2.2)$$

$$\varphi \zeta = 0, \quad \nu(\varphi E) = 0,$$
  
 $g(\varphi E, \varphi F) = g(E, F) + \nu(E)\nu(F),$ 

$$g(E, \zeta) = \nu(E), \qquad (2.3)$$
$$\varphi(E, F) = \varphi(F, E) = g(E, \varphi F)$$

for any vector fields  $E, F \in \chi(M)$ , where  $\chi(M)$  is the Lie algebra of vector fields on M. If  $\zeta$  is a killing vector field, the (para) contact structure is called a K-(para) contact. In such a case, we have

$$\nabla_E \zeta = \varphi E$$

Recently, the authors Haseeb and Prasad defined and studied the following notion:

**Definition 2.1.** A Lorentzian almost paracontact manifold M is called Lorentzian para-Kenmostu manifold if [11]

$$(\nabla_E \varphi)F = -g(\varphi E, F)\zeta - \nu(F)\varphi E$$

for any E, F on M.

In an LP-Kenmostu manifold, we have

$$\nabla_E \zeta = -E - \nu(E)\zeta, \qquad (2.4)$$

$$(\nabla_E \nu)F = -g(E,F) - \nu(E)\nu(F), \qquad (2.5)$$

where  $\nabla$  denotes the Levi-Civita connection respecting to the Lorentzian metric *g*. Furthermore, in an *LP*-Kenmotsu manifold, the following relations hold [11]:

$$g(R(E,F)G,\zeta) = \nu(R(E,F)G) = g(F,G)\nu(E) - g(E,G)\nu(F),$$

$$R(\zeta,E)F = -R(E,\zeta)F = g(E,F)\zeta - \nu(F)E,$$

$$R(E,F)\zeta = \nu(F)E - \nu(E)F,$$

$$R(\zeta,E)\zeta = E + \nu(E)\zeta,$$
(2.6)
$$S(E,\zeta) = (n-1)\nu(E) - S(\zeta,\zeta) = (n-1),$$

$$S(E,\zeta) = (n-1)\nu(E), \ S(\zeta,\zeta) = -(n-1),$$
 (2.7)

 $Q\zeta = (n-1)\zeta,$ 

for any  $E, F, G \in \chi(M)$ , where R, S and Q represent the curvature tensor, the Ricci tensor and the Q Ricci operator, respectively.

**Definition 2.2.** [19] An LP-Kenmotsu manifold M is said to be  $\nu$ -Einstein manifold if its  $S(\neq 0)$  is of the form

$$S(E, F) = ag(E, F) + b\nu(E)\nu(F),$$

where a and b are smooth functions on M. In particular, if b = 0, then M is termed as an Einstein manifold.

**Remark 2.1.** [12] In an LP-Kenmotsu manifold of n-dimension, S is of the form

$$S(E,F) = \left(\frac{r}{n-1} - 1\right)g(E,F) + \left(\frac{r}{n-1} - n\right)\nu(E)\nu(F),$$
(2.8)

where r is the scalar curvature of the manifold.

Lemma 2.1. In an n-dimensional LP-Kenmotsu manifold, we have

$$\zeta(r) = 2(r - n(n - 1)), \tag{2.9}$$

$$(\nabla_E Q)\zeta = QE - (n-1)E, \qquad (2.10)$$

$$(\nabla_{\zeta}Q)E = 2QE - 2(n-1)E,$$
 (2.11)

for any E on M.

Proof. Equation (2.8) yields

$$QE = (\frac{r}{n-1} - 1)E + (\frac{r}{n-1} - n)\nu(E)\zeta.$$
(2.12)

Taking the covariant derivative of (2.12) with respect to *F* and making use of (2.4) and (2.5), we lead to

$$(\nabla_F Q)E = \frac{F(r)}{n-1}(E + \nu(E)\zeta) - (\frac{r}{n-1} - n)(g(E, F)\zeta + \nu(E)F + 2\nu(E)\nu(F)\zeta).$$

By contracting F in the foregoing equation and using trace  $\{F \to (\nabla_F Q)E\} = \frac{1}{2}E(r)$ , we find

$$\frac{n-3}{2(n-1)}E(r) = \left\{\frac{\zeta(r)}{n-1} - (r - n(n-1))\right\}\nu(E),$$

which by replacing *E* by  $\zeta$  and using (2.1) gives (2.9). We refer the readers to see [13] for the proof of (2.10) and (2.11).

**Remark 2.2.** From the equation (2.9), it is noticed that if an n-dimensional LP-Kenmotsu manifold possesses the constant scalar curvature, then r = n(n - 1) and hence (2.8) reduces to S(E, F) = (n - 1)g(E, F). Thus the manifold under consideration is an Einstein manifold.

## 3. CRYS on LP-Kenmotsu Manifolds

Let the metric of an *n*-dimensional LP-Kenmotsu manifold be a conformal Ricci-Yamabe soliton, thus (1.2) holds. By differentiating (1.2) covariantly with respect to G, we have

$$(\nabla_G \mathcal{L}_K g)(E, F) = -2\sigma(\nabla_G S)(E, F) + \rho(Gr)g(E, F).$$
(3.1)

Since  $\nabla g = 0$ , then the following formula [18]

$$(\pounds_{\mathcal{K}}\nabla_{E}g - \nabla_{E}\pounds_{\mathcal{K}}g - \nabla_{[\mathcal{K},E]}g)(F,G) = -g((\pounds_{\mathcal{K}}\nabla)(E,F),G) - g((\pounds_{\mathcal{K}}\nabla)(E,G),F)$$

turns to

$$(\nabla_E \pounds_K g)(F,G) = g((\pounds_K \nabla)(E,F),G) + g((\pounds_K \nabla)(E,G),F).$$

Since the operator  $\pounds_{\mathcal{K}} \nabla$  is symmetric, therefore we have

$$2g((\pounds_{\mathcal{K}}\nabla)(E,F),G) = (\nabla_{E}\pounds_{\mathcal{K}}g)(F,G) + (\nabla_{F}\pounds_{\mathcal{K}}g)(E,G) - (\nabla_{G}\pounds_{\mathcal{K}}g)(E,F),$$

which by using (3.1) takes the form

$$2g((\pounds_{K}\nabla)(E,F),G) = -2\sigma[(\nabla_{E}S)(F,G) + (\nabla_{F}S)(G,E) - (\nabla_{G}S)(E,F)] + \rho[(Er)g(F,G) + (Fr)g(G,E) - (Gr)g(E,F)].$$
(3.2)

Putting  $F = \zeta$  in (3.2) and using (2.3), we find

$$2g((\pounds_{K}\nabla)(E,\zeta),G) = -2\sigma[(\nabla_{E}S)(\zeta,G) + (\nabla_{\zeta}S)(G,E) - (\nabla_{G}S)(E,\zeta)] +\rho[(Er)\nu(G) + 2(r - n(n-1))g(E,G) - (Gr)\nu(E)].$$
(3.3)

By virtue of (2.10) and (2.11), (3.3) leads to

$$2g((\pounds_{K}\nabla)(E,\zeta),G) = -4\sigma[S(E,G) - (n-1)g(E,G)] +\rho[(Er)\nu(G) + 2(r - n(n-1))g(E,G) - (Gr)\nu(E)].$$

By eliminating G from the foregoing equation, we have

$$2(\pounds_{\mathcal{K}}\nabla)(F,\zeta) = \rho g(Dr,F)\zeta - \rho(Dr)\nu(F) - 4\sigma QF$$

$$+[4\sigma(n-1) + 2\rho(r-n(n-1))]F.$$
(3.4)

If we take r as constant, then from (2.9) it follows that r = n(n-1), and hence (3.4) reduces to

$$(\pounds_{\mathcal{K}}\nabla)(F,\zeta) = -2\sigma QF + 2\sigma(n-1)F.$$
(3.5)

Taking covariant derivative of (3.5) with respect to E, we have

$$(\nabla_E \pounds_K \nabla)(F, \zeta) = (\pounds_K \nabla)(F, E) - 2\sigma\nu(E)[QF - (n-1)F]$$
(3.6)  
-  $2\sigma(\nabla_E Q)F.$ 

Again from [18], we have

$$(\pounds_{\mathcal{K}} R)(E, F)G = (\nabla_{E} \pounds_{\mathcal{K}} \nabla)(F, G) - (\nabla_{F} \pounds_{\mathcal{K}} \nabla)(E, G),$$

which by putting  $G = \zeta$  and using (3.6) takes the form

$$(\pounds_{\mathcal{K}} R)(E, F)\zeta = 2\sigma\nu(F)(QE - (n-1)E) - 2\sigma\nu(E)(QF - (n-1)F)$$
(3.7)  
$$-2\sigma((\nabla_{E}Q)F - (\nabla_{F}Q)E).$$

Putting  $F = \zeta$  in (3.7) then using (2.1), (2.2), (2.10) and (2.11), we arrive at

$$(\pounds_{\mathcal{K}}R)(E,\zeta)\zeta = 0. \tag{3.8}$$

The Lie derivative of (2.6) along K leads to

$$(\pounds_{\mathcal{K}}R)(E,\zeta)\zeta - g(E,\pounds_{\mathcal{K}}\zeta)\zeta + 2\nu(\pounds_{\mathcal{K}}\zeta)E = -(\pounds_{\mathcal{K}}\nu)(E)\zeta.$$
(3.9)

From (3.8) and (3.9), we have

$$(\pounds_{\mathcal{K}}\nu)(E)\zeta = -2\nu(\pounds_{\mathcal{K}}\zeta)E + g(E,\pounds_{\mathcal{K}}\zeta)\zeta.$$
(3.10)

Taking the Lie derivative of  $g(E, \zeta) = \nu(E)$ , we find

$$(\pounds_{\mathcal{K}}\nu)(E) = g(E, \pounds_{\mathcal{K}}\zeta) + (\pounds_{\mathcal{K}}g)(E, \zeta).$$
(3.11)

By putting  $F = \zeta$  in (1.2) and using (2.7), we have

$$(\pounds_{\mathcal{K}}g)(E,\zeta) = -\{2\sigma(n-1) + 2\Lambda - \rho n(n-1) - (p+\frac{2}{n})\}\nu(E),$$
(3.12)

where r = n(n-1) being used.

Taking the Lie derivative of  $g(\zeta, \zeta) = -1$  along K we lead to

$$(\pounds_{\mathcal{K}}g)(\zeta,\zeta) = -2\nu(\pounds_{\mathcal{K}}\zeta). \tag{3.13}$$

From (3.12) and (3.13), we find

$$\nu(\pounds_{\kappa}\zeta) = -\{\sigma(n-1) + \Lambda - \frac{\rho n(n-1)}{2} - \frac{1}{2}(p+\frac{2}{n})\}.$$
(3.14)

Now combining the equations (3.10), (3.11), (3.12) and (3.14), we find

$$\Lambda = \frac{\rho n(n-1)}{2} - \sigma (n-1) + \frac{1}{2}(p + \frac{2}{n}).$$
(3.15)

Thus we have

**Theorem 3.1.** Let (M, g) be an n-dimensional LP-Kenmotsu manifold admitting CRYS with constant scalar curvature tensor, then  $\Lambda = \frac{\rho n(n-1)}{2} - \sigma(n-1) + \frac{1}{2}(p + \frac{2}{n})$ .

**Corollary 3.1.** Let the metric of n-dimensional LP-Kenmotsu manifold is CRYS. Then we have

Values of $\sigma$ , $\rho$	Soliton type	Soliton constant	CRYS to be expanding,
			shrinking or steady
$\sigma=1,\  ho=0$	conformal Ricci soliton	$\Lambda = \frac{1}{2}(p + \frac{2}{n}) - (n - 1)$	CRYS is shrinking, steady
			and expanding if p >
			$\frac{2(n^2-n-1)}{n}$ , $p = \frac{2(n^2-n-1)}{n}$
			and $p < \frac{2(n^2 - n - 1)}{n}$ , resp.
$\sigma=$ 0, $ ho=$ 1	conformal Yam- abe soliton	$\Lambda = \frac{\frac{1}{2}(p + \frac{2}{n}) + \frac{n(n-1)}{2}}{\frac{n(n-1)}{2}}$	CRYS is shrinking, steady
			and expanding if p <
			$\frac{-(n^3-n^2+2)}{n}$ , $p = \frac{-(n^3-n^2+2)}{n}$
			and $p > \frac{-(n^3 - n^2 + 2)}{n}$ , resp.
$\sigma=1,\  ho=-1$	conformal Ein- stein soliton	$\Lambda = \frac{1}{2}(p + \frac{2}{n}) - \frac{(n-1)(n+2)}{2}$	CRYS is shrinking,
			steady and expand-
			$  ing if p < \frac{(n+1)(n^2-2)}{n},$
			$p = \frac{(n+1)(n^2-2)}{n}$ and
			$p > \frac{(n+1)(n^2-2)}{n}$ , resp.

## 4. Gradient CRYS on LP-Kenmotsu Manifolds

Let *M* be an *n*-dimensional *LP*-Kenmotsu manifold with g as a gradient CRYS. Then equation (1.3) can be written as

$$\nabla_E Dv + \sigma QE + (\Lambda - \frac{\rho r}{2} - \frac{1}{2}(\rho + \frac{2}{n}))E = 0, \qquad (4.1)$$

for all vector fields E on M, where D denotes the gradient operator of g. Taking the covariant derivative of (4.1) with respect to F, we have

$$\nabla_F \nabla_E D v = -\sigma\{(\nabla_F Q)E + Q(\nabla_F E)\} + \rho \frac{F(r)}{2}E$$

$$-(\Lambda - \frac{\rho r}{2} - \frac{1}{2}(\rho + \frac{2}{n}))\nabla_F E.$$

$$(4.2)$$

Interchanging E and F in (4.2), we lead to

$$\nabla_E \nabla_F D v = -\sigma \{ (\nabla_E Q)F + Q(\nabla_E F) \} + \rho \frac{E(r)}{2}F$$

$$-(\Lambda - \frac{\rho r}{2} - \frac{1}{2}(\rho + \frac{2}{n}))\nabla_E F.$$
(4.3)

By making use of (4.1)-(4.3), we find

$$R(E, F)Dv = \sigma\{(\nabla_F Q)E - (\nabla_E Q)F\} + \frac{\rho}{2}\{E(r)F - F(r)E\}.$$
(4.4)

Now from (2.8), we find

$$QE = (\frac{r}{n-1} - 1)E + (\frac{r}{n-1} - n)\nu(E)\zeta,$$

which on taking covariant derivative with repect to F leads to

$$(\nabla_{F}Q)E = \frac{F(r)}{n-1}(E+\nu(E)\zeta) - (\frac{r}{n-1}-n)(g(E,F)\zeta) + 2\nu(E)\nu(F)\zeta + \nu(E)F).$$
(4.5)

By using (4.5) in (4.4), we have

$$R(E,F)Dv = \frac{(n-1)\rho - 2\sigma}{2(n-1)} \{E(r)F - F(r)E\} + \frac{\sigma}{n-1} \{F(r)\nu(E)\zeta - E(r)\nu(F)\zeta\} - \sigma(\frac{r}{n-1} - n)(\nu(E)F - \nu(F)E).$$
(4.6)

Contracting forgoing equation along E gives

$$S(F, Dv) = -\left\{\frac{(n-1)^2 \rho - 2\sigma(n-2)}{2(n-1)}\right\}F(r) + \frac{\sigma(n-3)(r-n(n-1))}{n-1}\nu(F).$$
(4.7)

From the equation (2.8), we have

$$S(F, Dv) = \left(\frac{r}{n-1} - 1\right)F(v) + \left(\frac{r}{n-1} - n\right)\nu(F)\zeta(v).$$
(4.8)

Now by equating (4.7) and (4.8), then putting  $F = \zeta$  and using (2.1), (2.9), we find

$$\zeta(v) = \frac{r - n(n-1)}{n-1} \{ \sigma - (n-1)\rho \}.$$
(4.9)

Taking the inner product of (4.6) with  $\zeta$ , we get

$$F(v)\nu(E) - E(v)\nu(F) = \frac{\rho}{2} \{ E(r)\nu(F) - F(r)\nu(E) \},\$$

which by replacing E by  $\zeta$  then using (2.9) and (4.9), we infer

$$F(v) = -\frac{\sigma(r - n(n-1))}{n-1}\nu(F) - \frac{\rho}{2}F(r).$$
(4.10)

If we take r as constant, then from Remark 2.5, we get r = n(n-1). Thus (4.10) leads to F(v) = 0. This implies that v is constant. Thus the soliton under the consideration is trivial. Hence we state:

**Theorem 4.1.** If the metric of an n-dimensional LP-Kenmotsu manifold of constant scalar curvature tensor admitting a special type of vector field is gradient CRYS, then the soliton is trivial.

For v constant, (1.3) turns to

$$\sigma QE = -(\Lambda - \frac{\rho r}{2} - \frac{1}{2}(\rho + \frac{2}{n}))E$$

which leads to

$$S(E,F) = -\frac{1}{\sigma} \left(\Lambda - \frac{\rho n(n-1)}{2} - \frac{1}{2}(\rho + \frac{2}{n})\right) g(E,F), \quad \sigma \neq 0.$$
(4.11)

By putting  $E = F = \zeta$  in (4.11) and using (2.7), we obtain

$$\Lambda = \frac{\rho n(n-1)}{2} - \sigma(n-1) + \frac{1}{2}(p + \frac{2}{n}).$$
(4.12)

**Corollary 4.1.** If an n-dimensional LP-Kenmotsu manifold admits a gradient CRYS with the constant scalar curvature, then the manifold under the consideration is an Einstein manifold and  $\Lambda = \frac{\rho n(n-1)}{2} - \sigma(n-1) + \frac{1}{2}(p + \frac{2}{n})$ .

#### 5. Example

We consider the 5-dimensional manifold  $M^5 = \{(x_1, x_2, x_3, x_4, x_5) \in \mathbb{R}^5 : x_5 > 0\}$ , where  $(x_1, x_2, x_3, x_4, x_5)$  are the standard coordinates in  $\mathbb{R}^5$ . Let  $\varrho_1$ ,  $\varrho_2$ ,  $\varrho_3$ ,  $\varrho_4$  and  $\varrho_5$  be the vector fields on  $M^5$  given by

$$\varrho_1 = e^{x_5} \frac{\partial}{\partial x_1}, \ \varrho_2 = e^{x_5} \frac{\partial}{\partial x_2}, \ \varrho_3 = e^{x_5} \frac{\partial}{\partial x_3}, \ \varrho_4 = e^{x_5} \frac{\partial}{\partial x_4}, \ \varrho_5 = \frac{\partial}{\partial x_5} = \zeta,$$

which are linearly independent at each point of  $M^5$ . Let g be the Lorentzian metric defined by

$$g(\varrho_i, \varrho_i) = 1$$
, for  $1 \le i \le 4$  and  $g(\varrho_5, \varrho_5) = -1$ ,  
 $g(\varrho_i, \varrho_i) = 0$ , for  $i \ne j$ ,  $1 \le i, j \le 5$ .

Let  $\nu$  be the 1-form defined by  $\nu(E) = g(E, \rho_5) = g(E, \zeta)$  for all  $E \in \chi(M^5)$ , and let  $\varphi$  be the (1,1)-tensor field defined by

$$\varphi \varrho_1 = -\varrho_2, \ \varphi \varrho_2 = -\varrho_1, \ \varphi \varrho_3 = -\varrho_4, \ \varphi \varrho_4 = -\varrho_3, \ \varphi \varrho_5 = 0.$$

By applying linearity of  $\varphi$  and g, we have

$$u(\zeta) = g(\zeta, \zeta) = -1, \ \varphi^2 E = E + \nu(E)\zeta \text{ and } g(\varphi E, \varphi F) = g(E, F) + \nu(E)\nu(F)$$

for all  $E, F \in \chi(M^5)$ . Thus for  $\rho_5 = \zeta$ , the structure  $(\varphi, \zeta, \nu, g)$  defines a Lorentzian almost paracontact metric structure on  $M^5$ . Then we have

$$[\varrho_i, \varrho_j] = -\varrho_i,$$
 for  $1 \le i \le 4, j = 5,$   
 $[\varrho_i, \varrho_j] = 0,$  otherwise.

By using Koszul's formula, we can easily find we obtain

$$\nabla_{\varrho_i}\varrho_j = \begin{cases} -\varrho_5, & 1 \le i = j \le 4, \\ -\varrho_i, & 1 \le i \le 4, j = 5 \\ 0, & otherwise. \end{cases}$$

Also one can easily verify that

$$\nabla_E \zeta = -E - \eta(E)\zeta$$
 and  $(\nabla_E \varphi)F = -g(\varphi E, F)\zeta - \nu(F)\varphi E$ 

Therefore, the manifold is an *LP*-Kenmotsu manifold.

From the above results, we can easily obtain the non-vanishing components of R as follows:

$$\begin{aligned} R(\varrho_1, \varrho_2)\varrho_1 &= -\varrho_2, \ R(\varrho_1, \varrho_2)\varrho_2 = \varrho_1, \ R(\varrho_1, \varrho_3)\varrho_1 = -\varrho_3, \ R(\varrho_1, \varrho_3)\varrho_3 = \varrho_1, \\ R(\varrho_1, \varrho_4)\varrho_1 &= -v_4, \ R(\varrho_1, \varrho_4)\varrho_4 = \varrho_1, \ R(\varrho_1, \varrho_5)\varrho_1 = -\varrho_5, \ R(\varrho_1, \varrho_5)\varrho_5 = -\varrho_1, \\ R(\varrho_2, \varrho_3)\varrho_2 &= -\varrho_3, \ R(\varrho_2, \varrho_3)\varrho_3 = \varrho_2, \ R(\varrho_2, \varrho_4)\varrho_2 = -\varrho_4, \ R(\varrho_2, \varrho_4)\varrho_4 = \varrho_2, \\ R(\varrho_2, \varrho_5)\varrho_2 &= -\varrho_5, \ R(\varrho_2, \varrho_5)\varrho_5 = -\varrho_2, \ R(\varrho_3, \varrho_4)\varrho_3 = -\varrho_4, \ R(\varrho_3, \varrho_4)\varrho_4 = \varrho_3, \\ R(\varrho_3, \varrho_5)\varrho_3 &= -\varrho_5, \ R(\varrho_3, \varrho_5)\varrho_5 = -\varrho_3, \ R(\varrho_4, \varrho_5)\varrho_4 = -\varrho_5, \ R(\varrho_4, \varrho_5)\varrho_5 = -\varrho_4. \end{aligned}$$

Also, we calculate the Ricci tensors as follows:

$$S(\varrho_1, \varrho_1) = S(\varrho_2, \varrho_2) = S(\varrho_3, \varrho_3) = S(\varrho_4, \varrho_4) = 4, \quad S(\varrho_5, \varrho_5) = -4,$$

Therefore, we have

$$r = S(\varrho_1, \varrho_1) + S(\varrho_2, \varrho_2) + S(\varrho_3, \varrho_3) + S(\varrho_4, \varrho_4) - S(\varrho_5, \varrho_5) = 20.$$

Now by taking  $Dv = (\varrho_1 v)\varrho_1 + (\varrho_2 v)\varrho_2 + (\varrho_3 v)\varrho_3 + (\varrho_4 v)\varrho_4 + (\varrho_5 v)\varrho_5$ , we have

$$\nabla_{\varrho_1} Dv = (\varrho_1(\varrho_1 v) - (\varrho_5 v))\varrho_1 + (\varrho_1(\varrho_2 v))\varrho_2 + (\varrho_1(\varrho_3 v))\varrho_3 + (\varrho_1(\varrho_4 v))\varrho_4 + (\varrho_1(\varrho_5 v) - (\varrho_1 v))\varrho_5,$$

$$\begin{aligned} \nabla_{\varrho_2} Dv &= (\varrho_2(\varrho_1 v))\varrho_1 + (\varrho_2(\varrho_2 v) - (\varrho_5 v))\varrho_2 + (\varrho_2(\varrho_3 v))\varrho_3 + (\varrho_2(\varrho_4 v))\varrho_4 \\ &+ (\varrho_2(\varrho_5 v) - (\varrho_2 v))\varrho_5, \end{aligned}$$

$$\begin{aligned} \nabla_{\varrho_3} Dv &= (\varrho_3(\varrho_1 v))\varrho_1 + (\varrho_3(\varrho_2 v))\varrho_2 + (\varrho_3(\varrho_3 v) - (\varrho_5 v))\varrho_3 + (\varrho_3(\varrho_4 v))\varrho_4 \\ &+ (\varrho_3(\varrho_5 v) - (\varrho_3 v))\varrho_5, \end{aligned}$$

$$\begin{aligned} \nabla_{\varrho_4} Dv &= (\varrho_4(\varrho_1 v))\varrho_1 + (\varrho_4(\varrho_2 v))\varrho_2 + (\varrho_4(\varrho_3 v))\varrho_3 + (\varrho_4(\varrho_4 v) - (\varrho_5 v))\varrho_4 \\ &+ (\varrho_4(\varrho_5 v) - (\varrho_4 v))\varrho_5, \end{aligned}$$

$$\nabla_{\varrho_5} Dv = (\varrho_5(\varrho_1 v))\varrho_1 + (\varrho_5(\varrho_2 v))\varrho_2 + (\varrho_5(\varrho_3 v))\varrho_3 + (\varrho_5(\varrho_4 v))\varrho_4 + (\varrho_5(\varrho_5 v))\varrho_5.$$

Thus by virtue of (4.1), we obtain

$$\begin{cases} \varrho_{1}(\varrho_{1}v) - \varrho_{5}v = -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(\rho + \frac{2}{5})), \\ \varrho_{2}(\varrho_{2}v) - \varrho_{5}v = -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(\rho + \frac{2}{5})), \\ \varrho_{3}(\varrho_{3}v) - \varrho_{5}v = -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(\rho + \frac{2}{5})), \\ \varrho_{4}(\varrho_{4}v) - \varrho_{5}v = -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(\rho + \frac{2}{5})), \\ \varrho_{5}(\varrho_{5}v) = -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(\rho + \frac{2}{5})), \\ \varrho_{1}(\varrho_{2}v) = \varrho_{1}(\varrho_{3}v) = \varrho_{1}(\varrho_{4}v) = 0, \\ \varrho_{1}(\varrho_{2}v) = \varrho_{2}(\varrho_{3}v) = \varrho_{2}(\varrho_{4}v) = 0, \\ \varrho_{3}(\varrho_{1}v) = \varrho_{3}(\varrho_{2}v) = \varrho_{3}(\varrho_{4}v) = 0, \\ \varrho_{4}(\varrho_{1}v) = \varrho_{4}(\varrho_{2}v) = \varrho_{4}(\varrho_{3}v) = 0, \\ \varrho_{1}(\varrho_{5}v) - (\varrho_{1}v) = \varrho_{2}(\varrho_{5}v) - (\varrho_{2}v) = 0, \\ \varrho_{3}(\varrho_{5}v) - (\varrho_{3}v) = \varrho_{4}(\varrho_{5}v) - (\varrho_{4}v) = 0. \end{cases}$$

$$(5.1)$$

Thus the equations in (5.1) are respectively amounting to

$$\begin{aligned} e^{2x_5} \frac{\partial^2 v}{\partial x_1^2} - \frac{\partial v}{\partial x_5} &= -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(p + \frac{2}{5})), \\ e^{2x_5} \frac{\partial^2 v}{\partial x_2^2} - \frac{\partial v}{\partial x_5} &= -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(p + \frac{2}{5})), \\ e^{2x_5} \frac{\partial^2 v}{\partial x_3^2} - \frac{\partial v}{\partial x_5} &= -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(p + \frac{2}{5})), \\ e^{2x_5} \frac{\partial^2 v}{\partial x_4^2} - \frac{\partial v}{\partial x_5} &= -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(p + \frac{2}{5})), \\ \frac{\partial^2 v}{\partial x_5^2} &= -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(p + \frac{2}{5})), \\ \frac{\partial^2 v}{\partial x_5^2} &= -(\Lambda + 4\sigma - 10\rho - \frac{1}{2}(p + \frac{2}{5})), \\ \frac{\partial^2 v}{\partial x_1 \partial x_2} &= \frac{\partial^2 v}{\partial x_1 \partial x_3} &= \frac{\partial^2 v}{\partial x_1 \partial x_4} &= \frac{\partial^2 v}{\partial x_2 \partial x_3} &= \frac{\partial^2 v}{\partial x_2 \partial x_4} &= \frac{\partial^2 v}{\partial x_3 \partial x_4} = 0, \end{aligned}$$

$$e^{x_5}\frac{\partial^2 v}{\partial x_5 \partial x_1} + \frac{\partial v}{\partial x_1} = e^{x_5}\frac{\partial^2 v}{\partial x_5 \partial x_2} + \frac{\partial v}{\partial x_2} = e^{x_5}\frac{\partial^2 v}{\partial x_5 \partial x_3} + \frac{\partial v}{\partial x_3} = e^{x_5}\frac{\partial^2 v}{\partial x_5 \partial x_4} + \frac{\partial v}{\partial x_4} = 0.$$

From the above equations it is observed that v is constant for  $\Lambda = -4\sigma + 10\rho + \frac{1}{2}(p + \frac{2}{5})$ . Hence equation (4.1) is satisfied. Thus, g is a gradient RYS with the soliton vector field K = Dv, where v is constant and  $\Lambda = -4\sigma + 10\rho + \frac{1}{2}(p + \frac{2}{5})$ . Thus, Theorem 4.1 is verified.

**Conflicts of Interest:** The authors declare that there are no conflicts of interest regarding the publication of this paper.

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