A Statistical Study of Aftershock Sequences

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SUMMARY. — A comprehensive statistical study of the phenomenology of aftershock sequences is made in this paper. The spatial distribution of aftershocks indicates that they are mainly crustal events; however, deeper sequences also take place. The analysis of the distribution of aftershocks in 15 sequences with respect to time and magnitude leads to the statistical confirmation of a set of phenomenological laws describing the process, namely, the time-frequency law of hyperbolic decay of aftershock activity with time, the magnitude stability law, and the exponential magnitude-frequency distribution. The hypotheses involved are checked. The grouping of data and the statistical methods employed are chosen according to some basic well-confirmed assumptions regarding the nature of the process.

RIASSUNTO. — Questo articolo presenta un completo studio statistico della fenomenologia delle serie di repliche di terremoti. La distribuzione spaziale degli ipocentri delle repliche indica che si tratta di un processo che interessa principalmente la crosta terrestre; tuttavia non mancano esempi di serie di repliche profonde. L'analisi della distribuzione delle repliche in 15 serie rispetto al tempo e alla magnitudo conferma le leggi statistiche che descrivono lo sviluppo di una serie: la frequenza delle repliche decresce iperbolicamente nel tempo, la magnitudo è stabile nel tempo, e la distribuzione della magnitudo è esponenziale. Le ipotesi considerate sono sottoposte ad esame. La classificazione dei dati e i metodi statistici impiegati sono scelti in base ad alcune ipotesi che sono state ampiamente confermate.

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Introduction.

A comprehensive study of the phenomenology of aftershock sequences must include space, time, and magnitude distributions. Sequences which occurred in many parts of the world have been studied by various authors and the results are scattered in the geophysical literature. The determination of the phenomenological laws describing the aftershock process is a statistical problem, and it is therefore desirable to employ a consistent statistical procedure and to perform appropriate tests. We feel that some of the methods employed in the past were not rigorous; consequently, this paper is an attempt to present a unified procedure for the statistical study of aftershock sequences. For this purpose, 15 sequences have been analyzed in detail, even if some of them had already been studied according to different methods.

The study of the space distribution of aftershocks within a given sequence does not present particular problems. Its accuracy depends on the precision with which epicentral coordinates and focal depths are computed. It has been maintained (1) that aftershocks are essentially a shallow phenomenon. A review of available information partly confirms this view, but allowance must be made for notable exceptions.

The situation is more complicated with respect to time and magnitude distribution. The problem is basically that of finding a statistical relationship between the various quantities involved, and of estimating the parameters appearing in the statistical laws. First of all, when examining the data (which consist of the origin times t_i and of the individual magnitudes M_i of the aftershocks), one must be reasonably sure to be dealing with a complete set; that is, ideally no aftershock with $M \geq M^*$, where M^* is the minimum magnitude detected in a sequence, should be missing in the time interval considered. Moreover, it has been shown empirically by Suzuki (2) that the mode of grouping the data influences the results: therefore the mode of grouping should be as uniform as possible. Finally, it is desirable to employ a statistical procedure that does not contradict the underlying characteristics of the phenomenon observed, it is necessary to apply some statistical test to check the hypothesis being entertained, and confidence limits on the results must be given. More often than not, one or several of these conditions are not met in the study of

Table I - LIST OF AFTERSHOCK SEQUENCES SELECTED FOR DETAILED STUDY.

Sequence	Region	Reference	t_o	φ_o	λ_o	h_o	M_o	M^*	h	K
1. Alaska 1964	Alaska	Page (3)	Mar 28,03:36:13	61.00 N	147°.8 W	20 km	8.5	4.5	< 35 km	294
2. Aleutian I. 1957	Aleutian I.	Duda (4)	Mar 9,14:22:28	51.°3 N	175°.8 W	- 1	8.3	5.9	<150 »	205
3. Long Beach 1933	California	Benioff (5)	Mar 10,17:54:08	33°34′N	117°59′W	shallow	6.3	3.9	shallow	78
4. Desert Hot Sp. 1948))	Richter et al. (6)	Dec 4,15:43:17	33°56′N	116°23′W	shallow	6.5	3.0	$< 35 \; \mathrm{km}$	68
5. Kern County 1952	»	Richter (7)	Jul 21,11:52:14	35°.0′N	119°02′W	16 km	7.7	4.0	shallow	184
6. San Francisco 1957	»	Tocher (8)	Mar 22,19:44:21	37°40′N	122°29′W	shallow	5.3	2.0	<10 km	160
7. Salinas 1963	»	Udias (9)	Sept 14,19:46:17	36°52′N	121°38′W	3.3 km	5.4	1.0	<14 »	46
8. Parkfield 1966	»	McEvilly et al. (10)	Jun 28,04:26:13	35°57′N	120°29′W	shallow	5.5	2.0	<12 »	173
9. Chalkidike 1932	Greece	Papazachos et al. (11)	Sept 26,19:20:37	40°.5 N	23°.7 E	28 km	6.9	3.4	shallow	85
10. W. Thessaly 1954	»	»	Apr. 30,13:02:36	39°.3 N	22º 2 E	19 »	7.0	3.2	»	299
11. Amorgos 1956	»))	Jul 9,03:11:40	36°.7 N	25°.8 E	21 »	7.5	3.5	»	400
12. Magnesia 1957	_ »	»	Mar 8,12:21:13	39°.3 N	22°.6 E	18 »	6.8	3.0	»	291
13. Zante 1962	»	»	Apr 10,21:37:13	37°.6 N	20°.1 E	20 »	6.3	3.6	»	139
14. Cremasta 1966	»	Comninakis et al. (12)	Feb 5,02:01:43	39°.1 N	21°.6 E	20 »	5.9	3.4	»	103
15. Hawke's Bay 1931	New Zealand	Benioff (5)	Feb 3,10:15:00	39°.5 S	177°.0 E	_	7.6	4.1		71
										2

aftershock sequences, and the significance of the results is therefore debatable.

Among the several sequences for which origin times and magnitude of individual aftershocks have been published, 15 have been selected for detailed study. As far as possible we have tried to include sequences from different geographic regions, even if this implied considering a few sequences whose completeness may be in doubt. Since the determination of the statistical laws is more reliable when data are more abundant, no aftershock sequence consisting of less than 46 shocks has been included.

Table I lists the aftershock sequences whose time and magnitude distributions are studied in this paper; it includes region of occurrence, literature reference, main shock parameters (t_o origin time; ϕ_o , λ_o geographic coordinates; h_o focal depth; M_o magnitude), minimum magnitude of aftershock included (M^*), focal depth of aftershocks (h), and total number K of aftershocks with $M \geqslant M^*$ recorded in the first 100 days. (In sequence (1) the first day after the main shock is excluded from the count.) Focal depths are in kilometers; the term "shallow" is taken to mean "crustal", and often (especially in California) "upper crustal". Magnitude are given in the M-scale or M_L -scale, with the exception of sequence (1), where the m-scale has been employed.

It will be seen that the results of the analysis confirm the 3 basic laws describing the phenomenology of aftershock sequences, namely, the time-frequency law (Omori's law) as formulated by Mogi (¹³), the magnitude stability law (¹⁴), and the magnitude-frequency law (¹⁵).

SPATIAL DISTRIBUTION OF AFTERSHOCKS.

The spatial distribution of the shocks in an aftershock sequence is naturally related to the location of the main shock. The following considerations are based on a comprehensive survey of available data and are not limited to the sequences listed in Table I. If one traces on a map the boundary of the area in which the epicenters are located, the epicenter of the main shock is usually close to this boundary. Such is the case for all aftershock sequences of large earthquakes wich occurred in Japan from 1923 to 1963 (16). When aftershock activity takes place along a fault segment, as is frequently the case in California, the domain of the epicenters is approximately elliptical

with the long axis parallel to the active fault segment; often the main shock occupies, roughly speaking, one focus of the ellipse, and the aftershocks are concentrated toward the two ends. In some sequences the aftershock epicenters are clustered in a very small area (only a few kilometers in length and width), but this is rather exceptional; usually they are spread out over a much larger area. In the aftershock sequence of the Aleutian Islands earthquake of March 9, 1957 (4) the distribution of the epicenters follows closely the trend of tectonic activity along the Aleutian arc.

As to the focal depths of aftershocks, a review of available data by Page (1) indicates that, when hypocenter determinations are accurate, aftershocks are shallow events following a main shock which is itself shallow. Aftershock sequences therefore appear to be crustal phenomena, with the majority of shocks clustering in the upper layer of the crust ($h \leq 20$ km).

There are, however, some exceptions. In the region of Greece, the Southern Sporades earthquake of March 18, 1926 had a reported depth of 50 km and was followed by 18 aftershocks with $M \geq 3.9$ which were recorded in Athens; the sequence of the Anatolia earthquake of March 18, 1953 (h = 50 km) comprised 21 aftershocks in the first 13 days; the sequence of the Zante earthquake of November 15, 1959 (h = 55 km) consisted of 18 shocks; the focal depths of individual aftershocks, however, were not determined (11). In the aftershock sequence of the Kamchatka earthquake of November 4, 1952 the majority of shocks were located near the Mohorovičić discontinuity, but some of them had foci as deep as 60 km (17). The sequence following the Aleutian Islands earthquake of March 9, 1957 had an average focal depth of 74 km, and individual shocks were as much as 150 km deep (4).

Two more notable exceptions have occurred in Romania and in Central Asia. Iosif and Radu (18) have studied the aftershock sequence following an earthquake with M=7.4 and h=150 km that took place in the region of Vrancea, Romania, on November 10, 1940. The focal depths of aftershocks (3.3 $\leq M \leq 5.5$) were of the same order. Lukk (19) has studied the aftershock sequence of the Dzhurm earthquake of March 14, 1965, which occurred in the Pamir-Hindu Kush region and had a focal depth of 210 km. The observation period lasted for about 22 days, during which 390 aftershocks were recorded; their focal depths increased in time from 200 to more than 240 km.

The examples of recorded subcrustal aftershock sequences, however, form a very small part of the total number of sequences known to date; on the other hand, subcrustal earthquakes are themselves much less numerous than crustal ones. Therefore the comparative index of aftershock activity should be given by the ratio of the percentages of crustal and subcrustal earthquakes which are followed by a sequence. At present there is a bias due to instrumentation which favours the detection of shallow aftershock sequences whereas deep ones may go undetected. Thus, the conclusion that aftershocks are generally a shallow phenomenon has to be accepted and at the same time it must be realized that exceptions exist and that the data are far from complete.

TIME DISTRIBUTION OF AFTERSHOCKS.

It is customary to regard aftershocks as random events in time, whose frequency is governed by some time-decay law. Jeffreys (20), in a study of the aftershocks of the Tango, Japan, earthquake of March 7, 1927, found no sign of mutual dependence between aftershocks. That is, there was no indication that the chance of an aftershock in a given interval of time depended on anything but the time since the main shock, the aftershock frequency falling off with time according to Omori's law. The observed frequency showed only random departures from the law. It has since become a commonly accepted fact that aftershocks can be regarded as random independent events. It follows that any mathematical relationship relating time and frequency must not be interpreted as a physical "law" giving an exact correspondence, but as a statistical law of chance which is followed "on the average", observed frequencies showing random fluctuations from the theoretically expected values.

The fact that aftershock sequences consist of independent random events does not imply that they are a simple Poisson process. In a simple Poisson process the probability of occurrence of one event in a given time interval is constant for all t; this is obviously not the case for aftershocks, where the probability of occurrence depends on the time elapsed since the main shock. But, as Jeffreys (20) and many others have established, apart from the common dependence upon the main shock, no further mutual relation is found within the sequence.

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In this section the statistical decay law of aftershock activity is estimated for the 15 sequences listed in Table I. The data have been grouped according to a procedure suggested by Utsu (1). The origin time t_o of the main shock has been taken as origin of the time axis, $t_o = 0$. The origin times t_i of the aftershocks, obtained from the reference listed in Table I, have been expressed in terms of days after the main shock. The first day has been excluded from the analysis because of its possible incompleteness with respect to the number of shocks counted, due to the high frequency of aftershocks. Usually, aftershocks occurring in the time interval $1 \le t \le 100$ have been considered, unless a sequence comes to an end in a period of time shorter than 100 days. The time axis has been divided into logarithmically uniform intervals, such that their boundaries t_i^* satisfy the relation

$$\log t_i^* = 0.1 \ i \ , \ i = 0,1, \ldots, 20.$$

Now, if N_i is the number of aftershocks occurring in the time interval $\Delta t_i^* = t_{i+1}^* - t_i^*$, the quantity

$$n_i = \frac{N_i}{\Delta t_i^*}$$

represents the observed frequency per unit time interval. This observed frequency is associated with the centered value of the time interval concerned

$$t_i = \frac{t_i^* + t_{i+1}^*}{2}$$

so that one obtains a set of points (t_i, n_i) in the (t, n)-plane. The data arranged in this fashion are shown in Table II, the first column representing the centered time, the second the number of shocks in the time interval concerned, and the third the observed frequency. In the sequences(4), (7), (10), (12), (13) and (15), in which the number of shocks in some of the original time intervals was zero, the time intervals have been grouped two by two and n_i and t_i have been calculated accordingly.

The (t_i, n_i) -points usually show an approximately linear trend on doubly logarithmic paper. Consequently, it is reasonable to assume that the frequency of aftershocks per unit time n and the time t are related by an equation of the form

$$n(t) = a t^{-\beta}$$

Table II - Observed Frequency of Aftershocks.

t_i	N_{t}	n_4	t_i	N_4	n_i
(1) Alaska 19	64	2.837	2	3.075
	,		3.572	$ar{2}$	2.443
1.129	19	73.379	4.496	3	2.910
1.422	17	52.154		i	0.771
1.790	18	43.863	5.661	1	
2.254	16	30.970	7.126	•	0.612
2.837	19	29.213	8.972	2	0.972
$\frac{2.637}{3.572}$	13		11.295		0.386
		15.877	14.219	1	0.307
4.496	13	12.612	17.901	1	0.244
5.661	12	9.247	Charles Com		
7.126	19	11.630	RANGE TO SELECT		
8.972	11	5.348	(4) Desc	ert Hot Spr	ings 1948
11.295	10	3,862	Transfer in Advanced		- 0
14.219	17	5.215	1.292	3	5.129
17.901	22	5.361	2.048	2	2.158
22.536	19	3.678	3.246	4	2.723
28.371	6	0.923	5.145	7	3.006
35.717		1.588	8.155	8	2.168
	13		12.924	3	0.513
44.965	15	1.455		9	
56.607	13	1.002	20.484	5	0.539
71.264	9	0.551	32.465	5	0.340
89.716	13	0.632	51.453	1	0.043
			81.548	3	0.081
(2)	Aleutian I.	1957	and the		
1.129	6	23.172	(5) F	Kern County	y 1952
			1 120	5	10 310
1.422	9	27.611	1.129	5	19.310
$\substack{1.422\\1.790}$	9	$27.611 \\ 9.747$	1.422	7	21.475
$egin{array}{l} 1.422 \\ 1.790 \\ 2.254 \end{array}$	9 4 6	$27.611 \\ 9.747 \\ 11.614$	$1.422 \\ 1.790$	7 11	$21.475 \\ 26.805$
1.422 1.790 2.254 2.837	9 4 6 9	27.611 9.747 11.614 13.838	$egin{array}{c} 1.422 \ 1.790 \ 2.254 \end{array}$	7 11 13	21.475 26.805 25.163
1.422 1.790 2.254 2.837 3.572	9 4 6 9	27.611 9.747 11.614 13.838 10.992	$egin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \end{array}$	7 11 13 7	21.475 26.805 25.163 10.763
1.422 1.790 2.254 2.837 3.572 4.496	9 4 6 9 9	27.611 9.747 11.614 13.838 10.992 5.821	$egin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ \end{array}$	7 11 13 7 3	21.475 26.805 25.163 10.763 3.664
1.422 1.790 2.254 2.837 3.572	9 4 6 9	27.611 9.747 11.614 13.838 10.992	$egin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \end{array}$	7 11 13 7	21.475 26.805 25.163 10.763
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126	9 4 6 9 9 6 7 8	27.611 9.747 11.614 13.838 10.992 5.821	$egin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ \end{array}$	7 11 13 7 3 8 9	21.475 26.805 25.163 10.763 3.664
1.422 1.790 2.254 2.837 3.572 4.496 5.661	9 4 6 9 9 6 7	27.611 9.747 11.614 13.838 10.992 5.821 5.394	1.422 1.790 2.254 2.837 3.572 4.496	7 11 13 7 3 8	21.475 26.805 25.163 10.763 3.664 7.761
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126	9 4 6 9 9 6 7 8	27.611 9.747 11.614 13.838 10.992 5.821 5.394 4.897	$egin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ \end{array}$	7 11 13 7 3 8 9 4	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295	9 4 6 9 9 6 7 8 13	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \end{array}$	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972	7 11 13 7 3 8 9 4 7	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219	9 4 6 9 9 6 7 8 13 12	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \end{array}$	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295	7 11 13 7 3 8 9 4 7 8	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090
$\begin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ 8.972 \\ 11.295 \\ 14.219 \\ 17.901 \end{array}$	9 4 6 9 9 6 7 8 13 12 11	27.611 9.747 11.614 13.838 10.992 5.821 5.394 4.897 6.321 4.635 3.375 1.462	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219	7 11 13 7 3 8 9 4 7 8	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534
$\begin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ 8.972 \\ 11.295 \\ 14.219 \\ 17.901 \\ 22.536 \end{array}$	9 4 6 9 9 6 7 8 13 12 11 6	$egin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.83 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \\ 1.462 \\ 1.936 \\ \hline \end{array}$	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901	7 11 13 7 3 8 9 4 7 8 5	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218
$\begin{array}{c} 1.422\\ 1.790\\ 2.254\\ 2.837\\ 3.572\\ 4.496\\ 5.661\\ 7.126\\ 8.972\\ 11.295\\ 14.219\\ 17.901\\ 22.536\\ 28.371\\ \end{array}$	9 4 6 9 9 6 7 8 13 12 11 6	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \\ 1.462 \\ 1.936 \\ 1.076 \end{array}$	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536	7 11 13 7 3 8 9 4 7 8 5 5	21.475 26.805 25.163 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218
$\begin{array}{c} 1.422\\ 1.790\\ 2.254\\ 2.837\\ 3.572\\ 4.496\\ 5.661\\ 7.126\\ 8.972\\ 11.295\\ 14.219\\ 17.901\\ 22.536\\ 28.371\\ 35.717\\ \end{array}$	9 4 6 9 9 6 7 8 13 12 11 6 10 7	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \\ 1.462 \\ 1.936 \\ 1.076 \\ 1.466 \end{array}$	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371	7 11 13 7 3 8 9 4 7 8 5 5	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218 1.549 0.923
$\begin{array}{c} 1.422\\ 1.790\\ 2.254\\ 2.837\\ 3.572\\ 4.496\\ 5.661\\ 7.126\\ 8.972\\ 11.295\\ 14.219\\ 17.901\\ 22.536\\ 28.371\\ 35.717\\ 44.965\\ \end{array}$	9 4 6 9 9 6 7 8 13 12 11 6 10 7 12 6	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \\ 1.462 \\ 1.936 \\ 1.076 \\ 1.466 \\ 0.582 \\ \end{array}$	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717	7 11 13 7 3 8 9 4 7 8 5 5 8 6 6	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218 0.923 0.733
$\begin{array}{c} 1.422\\ 1.790\\ 2.254\\ 2.837\\ 3.572\\ 4.496\\ 5.661\\ 7.126\\ 8.972\\ 11.295\\ 14.219\\ 17.901\\ 22.536\\ 28.371\\ 35.717\\ 44.965\\ 56.607\\ \end{array}$	9 4 6 9 9 6 7 8 13 12 11 6 10 7 12 6 9	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \\ 1.462 \\ 1.936 \\ 1.076 \\ 1.466 \\ 0.582 \\ 0.694 \end{array}$	$\begin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ 8.972 \\ 11.295 \\ 14.219 \\ 17.901 \\ 22.536 \\ 28.371 \\ 35.717 \\ 44.965 \end{array}$	7 11 13 7 3 8 9 4 7 8 5 5 8 6 6	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218 1.549 0.923 0.733 0.485
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717 44.965 56.607 71.264	9 4 6 9 9 6 7 8 13 12 11 6 10 7 12 6	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \\ 1.462 \\ 1.936 \\ 1.076 \\ 1.466 \\ 0.582 \\ \end{array}$	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717	7 11 13 7 3 8 9 4 7 8 5 5 8 6 6	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218 0.923 0.733
$\begin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ 8.972 \\ 11.295 \\ 14.219 \\ 17.901 \\ 22.536 \\ 28.371 \\ 35.717 \\ 44.965 \\ 56.607 \end{array}$	9 4 6 9 9 6 7 8 13 12 11 6 10 7 12 6 9	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \\ 1.462 \\ 1.936 \\ 1.076 \\ 1.466 \\ 0.582 \\ 0.694 \end{array}$	$\begin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ 8.972 \\ 11.295 \\ 14.219 \\ 17.901 \\ 22.536 \\ 28.371 \\ 35.717 \\ 44.965 \\ 56.607 \\ 71.264 \end{array}$	7 11 13 7 3 8 9 4 7 8 5 5 8 6 6 6 5 3	$\begin{array}{c} 21.475 \\ 26.805 \\ 25.163 \\ 10.763 \\ 3.664 \\ 7.761 \\ 6.935 \\ 2.448 \\ 3.404 \\ 3.090 \\ 1.534 \\ 1.218 \\ 1.549 \\ 0.923 \\ 0.733 \\ 0.485 \\ 0.385 \\ 0.184 \\ \end{array}$
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717 44.965 56.607 71.264	9 4 6 9 9 6 7 8 13 12 11 6 10 7 12 6 9 3	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \\ 1.462 \\ 1.936 \\ 1.076 \\ 1.466 \\ 0.582 \\ 0.694 \\ 0.184 \end{array}$	$\begin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ 8.972 \\ 11.295 \\ 14.219 \\ 17.901 \\ 22.536 \\ 28.371 \\ 35.717 \\ 44.965 \\ 56.607 \end{array}$	7 11 13 7 3 8 9 4 7 8 5 5 8 6 6 6 5	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218 1.549 0.923 0.733 0.485 0.385
$\begin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ 8.972 \\ 11.295 \\ 14.219 \\ 17.901 \\ 22.536 \\ 28.371 \\ 35.717 \\ 44.965 \\ 56.607 \\ 71.264 \\ 89.716 \end{array}$	9 4 6 9 9 6 7 8 13 12 11 6 10 7 12 6 9 3	$\begin{array}{c} 27.611 \\ 9.747 \\ 11.614 \\ 13.838 \\ 10.992 \\ 5.821 \\ 5.394 \\ 4.897 \\ 6.321 \\ 4.635 \\ 3.375 \\ 1.462 \\ 1.936 \\ 1.076 \\ 1.466 \\ 0.582 \\ 0.694 \\ 0.184 \\ 0.340 \\ \end{array}$	$\begin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ 8.972 \\ 11.295 \\ 14.219 \\ 17.901 \\ 22.536 \\ 28.371 \\ 35.717 \\ 44.965 \\ 56.607 \\ 71.264 \\ 89.716 \\ \end{array}$	7 11 13 7 3 8 9 4 7 8 5 5 8 6 6 5 5 3 3	$\begin{array}{c} 21.475 \\ 26.805 \\ 25.163 \\ 10.763 \\ 3.664 \\ 7.761 \\ 6.935 \\ 2.448 \\ 3.404 \\ 3.090 \\ 1.534 \\ 1.218 \\ 1.549 \\ 0.923 \\ 0.733 \\ 0.485 \\ 0.385 \\ 0.184 \\ 0.146 \\ \end{array}$
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717 44.965 56.607 71.264 89.716	9 4 6 9 9 6 7 8 13 12 11 6 10 7 12 6 9 3 7	27.611 9.747 11.614 13.838 10.992 5.821 5.394 4.897 6.321 4.635 3.375 1.462 1.936 1.076 1.466 0.582 0.694 0.184 0.340	$\begin{array}{c} 1.422 \\ 1.790 \\ 2.254 \\ 2.837 \\ 3.572 \\ 4.496 \\ 5.661 \\ 7.126 \\ 8.972 \\ 11.295 \\ 14.219 \\ 17.901 \\ 22.536 \\ 28.371 \\ 35.717 \\ 44.965 \\ 56.607 \\ 71.264 \\ 89.716 \\ \end{array}$	7 11 13 7 3 8 9 4 7 8 5 5 8 6 6 6 5 3	$\begin{array}{c} 21.475 \\ 26.805 \\ 25.163 \\ 10.763 \\ 3.664 \\ 7.761 \\ 6.935 \\ 2.448 \\ 3.404 \\ 3.090 \\ 1.534 \\ 1.218 \\ 1.549 \\ 0.923 \\ 0.733 \\ 0.485 \\ 0.385 \\ 0.184 \\ 0.146 \\ \end{array}$
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717 44.965 56.607 71.264 89.716	9 4 6 9 9 8 13 12 11 6 10 7 12 6 9 3 7	27.611 9.747 11.614 13.838 10.992 5.821 5.394 4.897 6.321 4.635 3.375 1.462 1.936 1.076 1.466 0.582 0.694 0.184 0.340 1933	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717 44.965 56.607 71.264 89.716	7 11 13 7 3 8 9 4 7 8 5 5 5 8 6 6 6 5 3 3	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218 1.549 0.923 0.733 0.485 0.385 0.184 0.146
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717 44.965 56.607 71.264 89.716	9 4 6 9 9 6 7 8 13 12 11 6 10 7 12 6 9 3 7	27.611 9.747 11.614 13.838 10.992 5.821 5.394 4.897 6.321 4.635 3.375 1.462 1.936 1.076 1.466 0.582 0.694 0.184 0.340 1933 11.586 3.068	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717 44.965 56.607 71.264 89.716 (6) Se	7 11 13 7 3 8 9 4 7 8 5 5 5 8 6 6 6 5 3 3	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218 1.549 0.923 0.733 0.485 0.385 0.184 0.146 0 1957 27.034
1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717 44.965 56.607 71.264 89.716	9 4 6 9 9 8 13 12 11 6 10 7 12 6 9 3 7	27.611 9.747 11.614 13.838 10.992 5.821 5.394 4.897 6.321 4.635 3.375 1.462 1.936 1.076 1.466 0.582 0.694 0.184 0.340 1933	1.422 1.790 2.254 2.837 3.572 4.496 5.661 7.126 8.972 11.295 14.219 17.901 22.536 28.371 35.717 44.965 56.607 71.264 89.716	7 11 13 7 3 8 9 4 7 8 5 5 5 8 6 6 6 5 3 3	21.475 26.805 25.163 10.763 3.664 7.761 6.935 2.448 3.404 3.090 1.534 1.218 1.549 0.923 0.733 0.485 0.385 0.184 0.146

Table II - Continued

t_i	N_i	n_t	t_i	N_i	n_i
2.254	6	11.614	(9)	Chalkidike	1932
2.837	2	3.075	()		
3.572	5	6.107	1.790	6	14.621
4.496	6	5.821	2.254	4	7.743
5.661	2	1.541	2.837	4	6.150
7.126	$\overline{2}$	1.224	3.572	12	14.656
8.972	3	1.459	4.496	9	8.731
11.295	$\frac{0}{2}$	0.772	5.661	6	4.624
14.219	$\frac{2}{2}$	0.614	7.126	6	3.673
	6				
17.901		1.462	8.972	6	2.917
22.536	5	0.968	11.295	6	2.317
28.371	3	0.451	14.219	8	2.454
35.717	3	0.366	17.901	4	0.975
44.965	4	0.388	22,536	1	0.194
56.607	4	0.308	28.371	1	0.154
71.264		0.061	35.717	2	0.244
89.716	6	0.292	44.965	The second	0.097
			56.607	2	0.154
			71.264	1 3 4	0.061
	(7) Salinas 1963		89.716	4	0.195
1.292	3	5.129			
2.048	2	2.158	(10) W	estern These	saly 1954
3.246	4	$\frac{2.138}{2.723}$			
	· · · · · · · · · · · · · · · · · · ·		1.292	13	22.226
5.145	6	2.577	2.048	9	9.709
8.155	3	0.813	3.246	13	8.849
12.924	8	1.368	5.145	24	10.307
20.484	2	0.216	8.155	32	8.671
			12.924	20	3.419
			20.484	11	1.187
	(8) Parkfield 196	6	32,465	34	2.314
			51.453	26	1.117
1.129	4	15.448	81.548	7	0.190
1.422	7	21.475	01.040	,	0.190
1.790	5	12.184			
2.254	10	19.356	(1)	1) Amorgos	1956
2.837	2	3.075	(1.	, 111101508	
3.572	4	4.885	1.129	27	104.275
4.496	9	8.731	1.422	27	82.832
5.661	9 5	$\frac{8.731}{3.853}$	1.790	26	63.358
			2.254	24	46.455
7.126	3	1.836	2.837	6	9.225
8.972	3	1.459	3.572	12	14.656
11.295	3	1.159		. –	
14.219	5	1.534	4.496	13	12.612
17.901	6	1.462	5.661	5	3.853
22.536	5	0.968	7.126	10	6.121
28.371	9	1.384	8.972	4	1.945
35.717	10	1.221			
44.965	5	0.485	/10	2) Magnesia	1057
56.607	4	0.308	(12	.) Magnesia	1001
71.264	4	0.245	1.292	13	22.226
89.716	5	0.243	2.048	4	4.31

Table II - Continued

t_i	N_i	n^{j}	t_i	N_i	n_i
3.246	23	15.655	5.661	6	4.624
5.145	24	10.307	7.126	2	1,224
8.155	24	6.503	8.972	uper launia	0.486
12,924	13	2.223	11.295	6	2.317
20.484	15	1.618	14.219	4	1.227
32.465	6	0.408	17.901	$\frac{2}{7}$	0.487
51.453	4	0.172	22.536	7	1,355
81.548	38	1.030	28.371		0.154
			35.717	14	1.710
			44.965	4	0.388
	(13) Zante 1962		56.607	7	0.539
	(/		71.264	6	0.367
1.292	11	18.807	89.716	18	0.875
2.048	7	7.551			
3.246	6	4.084	(1-)	n 1.1 n	1001
5.145	11	4.724	(15)	Hawke's Bay	C 1931
8.155	13	3.523	1.292	2	3.419
12.924	3	0.513	2.048	4	4.315
20.484	19	2.050	3.246	3	2.042
32.465	5	0.340	5.145	9	3.865
			8.155	5	1.355
			12.924	16	$\frac{1.336}{2.736}$
(14) Cremasta 196	16	20.484	7	0.755
,	, 0.0		32.465	4	$0.755 \\ 0.272$
3.572		1.221	51.453	5	0.215
4.496	4	3.881	81.548	3	0.081

that is, Omori's law. The commonest procedure for estimating the parameters a and β is the least squares method, which has been applied to the great majority of aftershock sequences whose time distribution has been investigated so far. Accordingly, relation [1] is linearized by taking logarithms on both sides

$$\log n(t) = \log a - \beta \log t.$$
 [2]

Then, setting

$$\log n(t) = y, \log t = x, \log a = a^*, -\beta = \beta^*$$

the following model is obtained

$$y = a^* + \beta_x^*;$$

that is, the expected value of y is a linear function of x. Therefore

$$y_i = a^* + \beta^* x_i, + \varepsilon_i$$

i.e., the observation y_i consists of the expected value at the given x_i plus a random fluctuation ε_i .

In the method of least squares, the parameters a^* and p^* are chosen in such a way that the sum of squares of the vertical distances of the points from the regression line is minimized. For the linear model, this sum of squares is

$$\xi(a^*, \beta^*) = \sum_{i=1}^k \varepsilon_i^2 = \sum_{i=1}^k (y_i - a^* - \beta^* x_i)^2$$

and the necessary conditions for ξ (a^* , β^*) to be minimum are

$$\frac{\partial \xi}{\partial a^*} = 0 \quad , \quad \frac{\partial \xi}{\partial \beta^*} = 0$$

from which one obtains

$$\beta^* = \frac{k \sum_{i=1}^k x_i \ y_i - \sum_{i=1}^k x_i \sum_{i=1}^k y_i}{k \sum_{i=1}^k x_i^2 - \left(\sum_{i=1}^k x_i\right)^2}$$

$$a^* = \bar{y} - \bar{\rho}^* \bar{x}, \ \bar{y} = \frac{1}{k} \sum_{i=1}^k y_i, \ \bar{x} = \frac{1}{k} \sum_{i=1}^k x_i$$
[3]

According to the Gauss-Markov theorem (22), the estimates of the parameters calculated according to the least squares method will be unbiased and of maximum efficiency if, and only if, the linear hypothesis is such that the random fluctuation ε has zero mean and constant variance (independent of x). In other words, if the calculated regression line is to give the expected value of y for each x, the observed values must be uncorrelated, and the probability distribution of y for each x must be symmetric. In many cases, especially when the fluctuation ε can be considered to be the sum of many independent factors, the conditions of the Gauss-Markov theorem are satisfied and the probability distribution of y may be regarded as approximately normal for every x. In other cases, however, the matter is very debatable, particularly when a transformation of coordinates is performed in order to linearize the least square model.

It turns out that, under reasonable assumptions, neither the original model [1] nor the linearized model [2] satisfy the conditions of the Gauss-Markov theorem. This has been noted, for instance, by Page (3). An appropriate method must take into account the prob-

ability distribution of n for each interval of time considered. On the other hand, it has been suggested (-3) that the least squares method is approximately correct also when the conditions for its theoretical validity are not realized in practice. In order to clarify these matters, in the sequel we shall estimate the parameters in equation [1] by means of both the least squares method for the linearized model [2], and the maximum likelihood method (24), which takes into account the distribution of n for each time interval. First of all, we analyze this distribution.

In the following discussion it is assumed that the deviations of the observed values n_i from the expected value n(t) reflect actual random fluctuations and that other contributions to such deviations are negligible. Since the only other sources of deviations are errors of measurement, and these have been reduced to a minimum by excluding the first day after the main shock and by counting only well-defined aftershocks with $M \geq M^*$, the assumption is most probably correct. It is also assumed that, if the number of shocks expected in the *i-th* interval is $E(N_i)$, the probability that the observed number is N_i , is given by

$$p(N_t; E(N_t)) = \frac{[E(N_t)]^{N_t}}{N_t!} e^{-E(N_t)}$$
[4]

i.e., the number of shocks in each time interval is given by a Poisson distribution. The Poisson distribution is the most fundamental distribution for such discrete variates as the number of shocks in a given time interval, and it has been postulated for the case of aftershocks by a number of authors, e.g., recently, by Utsu (21) and Page (3).

The expected value $E(N_i)$ in the *i-th* interval is given by

$$E(N_i) = \int_{t_i^*}^{t_{i+1}^*} n(t) dt = \int_{t_i^*}^{t_{i+1}^*} a t^{-\beta} dt \simeq a t_i^{-\beta} \Delta t_i^*$$
 [5]

where the approximation is introduced in order to avoid using the integral in equation [4]; this is necessary because β is unknown and could be unity. The approximation has been checked numerically for some randomly selected samples and the error was found to be negligible. Then relation [4] becomes

$$p(N_i; a, \beta) = \frac{(a t_i^{-\beta} \Delta t_i^*)^{N_i}}{\bar{N}_i!} e^{-a t_i^{-\beta} \Delta t_i^*}$$
[6]

The principle of the method of maximum likelihood is to take estimates of the unknown parameters that maximize the probability of obtaining the observed sample. Considering a sample of k independent values, each with a probability distribution $p(N_i; a, \beta)$, the probability that the sample consists precisely of these k values is

$$L(\alpha, \beta) = \prod_{t=1}^{k} p(N_t; \alpha, \beta)$$
 [7]

The function $L(a, \beta)$ is called the likelihood function. The necessary condition for $L(a, \beta)$ to have a maximum is

$$\frac{\partial L}{\partial a} = 0 \quad , \quad \frac{\partial L}{\partial \beta} = 0 \, .$$

Since $\ln L(a,\beta)$ (where \ln stands for the natural logarithm) attains its maximum for the same values of a and β as $L(a,\beta)$ itself, it is the function $\ln L(a,\beta)$ which is commonly maximized. It follows from equations [6] and [7] that the likelihood function in the present problem is

$$L(\alpha,\beta) = \prod_{i=1}^{k} \frac{(\alpha t_i^{-\beta} \Delta t_i^*)^{N_i}}{N_i!} e^{-\alpha t_i^{-\beta} \Delta t_i^*}$$

i.e.,

$$\ln L(a,\beta) = \sum_{i=1}^k \ln \left[\frac{a \, t_i^{-\beta} \, \Delta \, t_i^*)^{N_i}}{N_i!} \, e^{-a \, t_i^{-\beta} \, \Delta \, t_i^*} \right]$$

from which one obtains

$$\ln L(\alpha, \beta) = \ln \alpha \sum_{i=1}^k N_i - \beta \sum_{i=1}^k N_i \ln t_i + \sum_{i=1}^k N_i \ln \Delta t_i^* - \frac{1}{2} \sum_{i=1}^k t_i^{-\beta} \Delta t_i^* - \sum_{i=1}^k \ln N_i!$$

The maximum likelihood estimates of α and β are obtained by solving the equations

$$\begin{cases}
\frac{\partial \ln L(\alpha, \beta)}{\partial \alpha} = \frac{1}{\alpha} \sum_{i=1}^{k} N_i - \sum_{i=1}^{k} t_i^{-\beta} \Delta t_i^* = 0 \\
\frac{\partial \ln L(\alpha, \beta)}{\partial \beta} = -\sum_{i=1}^{k} N_i \ln t_i + \alpha \sum_{i=1}^{k} t_i^{-\beta} \Delta t_i^* \ln t_i = 0
\end{cases}$$

i.e.,

$$\begin{cases} \sum_{i=1}^{k} N_{i} - a \sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*} = 0 \\ \sum_{i=1}^{k} N_{i} \ln t_{i} - a \sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*} \ln t_{i} = 0 \end{cases}$$
[8]

Equations [8] are the normal equations in the maximum likelihood method. From them one obtains,

$$\alpha = \frac{\sum_{i=1}^{k} N_{i}}{\sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*}} = \frac{\sum_{i=1}^{k} N_{i} \ln t_{i}}{\sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*} \ln t_{i}}$$
[9]

i.e.,

$$F(\beta) = \sum_{i=1}^{k} N_i \sum_{i=1}^{k} t_i^{-\beta} \Delta t_i^* \ln t_i - \sum_{i=1}^{k} N_i \ln t_i \sum_{i=1}^{k} t_i^{-\beta} \Delta t_i^* = 0. \quad [10]$$

Equation [10] must be solved for β , and then α can be obtained from [9].

A program has been written to solve equation [10] by the secant method (25). The first approximation to the root of [10] and to the estimate of α has been obtained by means of the least squares method for the linearized model, according to formulas [3]. The estimates α_1 , β_1 obtained by the least squares method, when compared with the maximum likelihood estimates, give an idea on how statistically reliable the least squares method is when the underlying assumptions are not met.

The numerical results for the 15 sequences are summarized in Table III. From left to right, the columns indicate the sequence involved, the least square estimates a_1, β_1 , and the maximum likelihood estimates a, β . The parameter which characterizes a sequence is the decay parameter β , which measures the rate of decay in time of the frequency of aftershocks. It can be seen that differences between β_1 and β are present but not very large. The decay parameter is usually around unity.

Now, we proceed to check the validity of the time-frequency law. If the expected value of the frequency n varies in time according to equation [1], and using the approximation expressed by [5], the mean and the variance of the Poisson-distributed number of shocks are in each interval

$$E(N) = V(N) = n(t) \Delta t^* = a t^{-\beta} \Delta t^*$$

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and therefore the standard deviation of n is

$$D(n) = \sqrt{\frac{\alpha t^{-\beta}}{\Delta t^*}}.$$

Table III ESTIMATES OF THE PARAMETERS IN THE TIME-FREQUENCY LAW

Sequence	Least squar	e estimates	Maximum likelihood estimates		
	a_1	eta_1	а	β	
(1) Alaska	76.824	1.097	77.547	1.085	
(2) Aleutian I.	32.611	0.997	33.662	0.984	
(3) Long Beach	10.956	1.302	11.731	1.301	
(4) Desert Hot Sp.	9.303	1.075	9,970	1.042	
(5) Kern County	36.484	1.150	39.047	1.154	
(6) San Francisco	19.070	1.103	22.720	1.120	
(7) Salinas	6.651	0.902	6.778	0.844	
(8) Parkfield	20.622	0.981	23.424	0.998	
(9) Chalkidike	47.918	1.448	45.392	1.355	
(10) W. Thessaly	32.600	0.949	34.211	0.917	
(11) Amorgos	150.139	1.882	151.694	1.830	
(12) Magnesia	27.607	0.988	29.761	0.918	
(13) Zante	20.176	1.063	19.565	0.970	
(14) Cremasta	4.210	0.518	4.738	0.481	
(15) Hawke's Bay	8.639	0.907	11.135	0.941	

A general theorem which holds for an arbitrary distribution with a second moment is Tchebychev's theorem (24). It states that, if X is a random variable with mean E(X) and standard deviation D(X), then the following inequality holds

$$P\left\{\mid X - E(X) \mid \geqslant k \ D(X)\right\} \leqslant \frac{1}{k^2}$$

where k is an arbitrary positive number. In other words, the probability that X assumes values outside the interval $E(X) \pm k D(X)$ is less than $1/k^2$. Conversely, from the viewpoint of sampling, in the

long run less than $(100/k^2)\%$ of the values assumed by X will fall outside the interval.

The application of Tchebychev's inequality to check the validity of the time-frequency law has two limits, namely, it yields rather weak conditions, and the sample size is small. Nevertheless it gives a reasonably safe criterion for rejecting the validity of the law for sequences that show too wide fluctuations. Regarding the observed values of n as the result of random sampling from a population whose expected value varies with time according to [1], and choosing k=2, Tchebychev's inequality takes the form

$$P\Big\{|n-\alpha t^{-\beta}|\geqslant 2 \ \sqrt{\frac{\alpha t^{-\beta}}{\Delta t^*}}\Big\}\leqslant \frac{1}{4}$$

that is, in the long run less than 25% of the observed values n_t should fall outside the interval $E(n) \pm 2 D(n)$. If this condition is not satisfied, the assumption regarding the variation with time of the expected value n(t) must be rejected.

In order to determine a confidence band according to Tchebychev's inequality, therefore, the quantities E(n) + 2 D(n), E(n) - 2 D(n)have been computed at selected t_i , i = 1, 2, ..., k, and compared with the observed n_i . Table IV gives the results, showing in the columns from left to right the sequence, the total number of data points, the number of points outside the confidence band, and their percentage. Consequently, according to the selected criterion, sequences (10), (12), (13) and (14) do not follow the assumed timefrequency law, inasmuch as the observed frequency cannot be explained only in terms of random fluctuations from the law. The other 11 sequences appear to follow the law and the fit is generally fairly good. All the sequences which show considerable departures from the assumed law have occurred in the region of Greece; this fact might have some geotectonic significance. However, there exists the possibility that relatively poor instrumentation plays a part in some apparent irregularities. Sequence (14) originated under peculiar conditions; the frequency of aftershocks in it appears to be correlated to the variations in the water loading of a nearby artificial lake (12).

Figures 1 to 5 display on doubly logarithmic paper the results for some of the 11 sequences which appear to decay according to the postulated law. The dots represent the data points, the full line the fitted n(t), and the broken lines the confidence limits.

Table IV - FLUCTUATIONS OF THE OBSERVED FREQUENCIES IN THE DECAY OF AFTERSHOCK ACTIVITY WITH TIME.

Sequence	Total number of points	Points outside confidence band	% Points outside
(1) Alaska	20	2	10 %
(2) Aleutian I.	20	0	
(3) Long Beach	13	0	
(4) Desert Hot Sp.	10	0	union <u>dia</u> raban
(5) Kern County	20	0	The second
(6) San Francisco	20	0	
(7) Salinas	7	0	<u> </u>
(8) Parkfield	20	0	
(9) Chalkidike	18	1	5.6%
(10) W. Thessaly	10	3	30 %
(11) Amorgos	10	tivat ler per	10 %
(12) Magnesia	10	7	70 %
(13) Zante	8	2	25 %
(14) Cremasta	15	4	26.7%
(15) Hawke's Bay	10	1	10 07

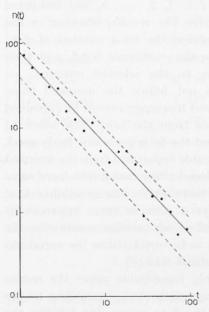


Fig. 1. – Time distribution of aftershocks: Alaska 1964.

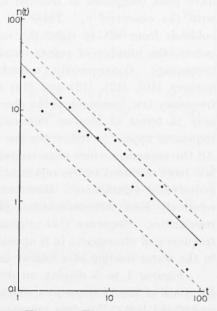


Fig. 2. – Time distribution of aftershocks: Aleutian Islands 1957.

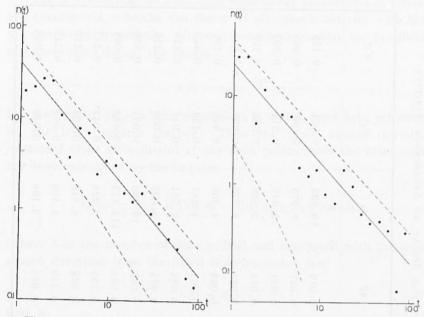


Fig. 3. – Time distribution of aftershocks: Kern County 1952.

Fig. 4. - Time distribution of aftershocks: San Francisco 1957.

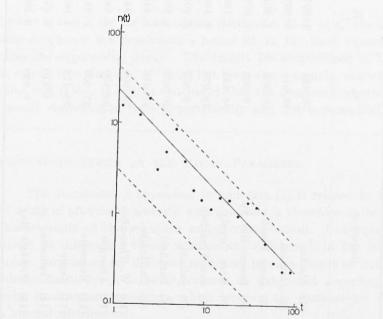


Fig. 5. - Time distribution of aftershocks: Parkfield 1966.

Table V - Comparison of Hyperbolic and Exponential Decay of Aftershock Activity.

Sequence	a_1	eta_1	8 e ²	a ₂	eta_2	8'e ²
(1) Alaska	76.824	1.097	0.016	18.295	0.051	0.123
(2) Aleutian I.	32.611	0.997	0.027	9.575	0.050	0.069
(3) Long Beach	10.956	1.302	0.031	5.917	0.211	0.052
(4) Desert Hot Sp.	9.303	1.075	0.078	2.529	0.055	0.109
(5) Kern County	36.484	1.150	0.025	8.570	0.056	0.105
(6) San Francisco	19.070	1.103	0.063	4.308	0.049	0.197
(7) Salinas	6.651	0.902	0.061	4.281	0.136	0.040
(8) Parkfield	20.622	0.981	0.038	5.751	0.046	0.119
(9) Chalkidike	47.918	1.448	0.075	5.093	0.060	0.216
(10) W. Thessaly	32.600	0.949	0.063	10.796	0.051	0.058
(11) Amorgos	150.139	1.882	0.035	112.473	0.482	0.062
(12) Magnesia	27.607	0.988	0.133	7.264	0.044	0.247
(13) Zante	20.176	1.063	0.076	8.346	0.103	0.116
(14) Cremasta	4.210	0.518	0.119	1.468	0.016	0.142
(15) Hawke's Bay	8.639	0.907	0.073	3.108	0.050	0.046

Its is interesting to note that most decay phenomena in physics are exponential, whereas the decay of aftershock activity with time appears to be hyperbolic. In order to check rapidly the possibility of exponential decay, the model

$$n(t) = a_2 e^{-\beta_2 t}$$

has been linearized, and the parameters α_2 and β_2 have been estimated by the least squares method. Then the mean square deviation (standard error of estimate) of the data points from the fitted curve has been calculated by the formula

$$s_e^{-2} = \frac{1}{k-2} \sum_{i=1}^{k} [\log n_i - (\log a_2 - \beta_2 t_i \log e)]^2$$

(where k is the number of data points) and compared with the mean square deviation from the fitted time-frequency law

$$n(t) = \alpha_1 t^{-\beta_1}$$

that is,

$$s_{\epsilon}^2 = \frac{1}{k-2} \sum_{i=1}^k [\log n_i - (\log \alpha_i - \beta_i \log t_i)]^2$$

(where a_1 and β , are the least square estimates). If $s_e^2 < s_e'$ the assumed time-frequency law represents a better fit, in the least square sense, than the exponential decay. The results are summarized in Table V. It can be seen that $s_e^2 < s_e^{2'}$ in all but three cases, namely, sequences (7), (10), and (15). It can be concluded that the frequency of aftershocks usually decreases in time hyperbolically and not exponentially.

CONFIDENCE LIMITS ON THE DECAY PARAMETER.

The parameter β appearing in equation [1] is related to the rate of decay of aftershock activity with time and is therefore an important characteristic of the sequence under consideration. Consequently, it would be interesting to see whether the differences in the computed decay parameters for different sequences are significant or not. Confidence limits on β , however, cannot be calculated according to the usual least square procedure, which assumes the fluctuations to have a normal distribution.

An approximate procedure for calculating confidence limits on β can be based on the addition theorem for the Poisson distribution (²⁴). Since in this procedure the parameter α is assumed to be known exactly, it will yield no more than an indication of the range in which the real value of β is likely to fall.

The addition theorem for the Poisson distribution states that, if N_i , i = 1, 2, ..., k, is a sequence of stochastically independent and Poisson-distributed random variables with expected values $E_i = E(N_i)$, then the sum $X = \sum_{i=1}^k N_i$ will be Poisson-distributed with expected value $E(X) = \sum_{i=1}^k E_i$. In the present case, N_i being the number of shocks in the *i-th* time interval, the expected values are

$$E_i = a t_i^{-\beta} \Delta t_i^*$$

and

$$E(X) = \alpha \sum_{i=1}^{k} t_{i}^{-\rho} \Delta t_{i}^{*}.$$

The value of E(X) for each of the 11 sequences which follow the time-frequency law is such that the distribution of X can be approximated by the normal distribution. Therefore the standardized variable

$$X^* = \frac{X - E(X)}{D(X)}$$

is approximately normally distributed with $E(X^*) = 0$, $V(X^*) = 1$. Therefore, the probability that X^* assumes a value in the interval (λ_1, λ_2) is

$$P\left\{\lambda_1 \leqslant X^* \leqslant \lambda_2\right\} = \Phi(\lambda_2) - \Phi(\lambda_1)$$

where $\Phi(X^*)$ is the normal distribution function. In particular

$$P\left\{\,-\,1.96\leqslant X^*\leqslant 1.96\,\right\}=0.95\,.$$

Then the approximate 95% confidence limits on the decay parameter are obtained by solving for β the equations

$$X^* = \pm 1.96$$

that is, recalling the definition of X^* ,

$$\frac{\sum_{i=1}^{k} N_{i} - a \sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*}}{\sqrt{a \sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*}}} = \pm 1.96.$$

Naturally, the central value of X^* is obtained when α , β are given by the maximum likelihood estimates. Assuming the value of α to be known exactly, the above equations take the form

$$\begin{cases} \sum_{i=1}^{k} N_{i} - a \sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*} - 1.96 & \sqrt{a \sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*}} = 0 \\ \sum_{i=1}^{k} N_{i} - a \sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*} + 1.96 & \sqrt{a \sum_{i=1}^{k} t_{i}^{-\beta} \Delta t_{i}^{*}} = 0 \end{cases}$$
[11]

Equations [11] have been solved by the secant method. The results for the 11 sequences with time-frequency law of the form given by [1] are shown in Table VI. It can be seen that the 95% confidence limits on the maximum likelihood estimate always contain the least square estimate of β , which therefore appears to be a good approximation to the real value of the decay parameter.

Table VI - Confidence limits on the decay parameter.

Sequences	95% Confidence limits			
(1) Alaska	$1.033 \leqslant \beta \leqslant 1.139$			
(2) Aleutian I.	$0.920 \leqslant \beta \leqslant 1.052$			
(3) Long Beach	$1.009 \leqslant \beta \leqslant 1.650$			
(4) Desert Hot Sp.	$0.913 \leqslant \beta \leqslant 1.185$			
(5) Kern County	$1.072 \leqslant \beta < 1.242$			
(6) San Francisco	$1.020 \leqslant \beta \leqslant 1.229$			
(7) Salinas	$0.645 \leqslant \beta \leqslant 1.064$			
(8) Parkfield	$0.919 \leqslant \beta \leqslant 1.082$			
(9) Chalkidike	$1.254 \leqslant \beta \leqslant 1.464$			
(11) Amorgos	$1.644 \leqslant \beta \leqslant 2.032$			
(15) Hawke's Bay	$0.829 \leqslant \beta \leqslant 1.051$			

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The values listed in Table VI must not be considered as exact. The limits rounded to the second decimal digit are probably fairly reliable. Most β -values cluster around the 0.9-1.2 range. Sequence (11) stands by itself, showing a very rapid decay in activity. Also sequence (9) has a decay coefficient somewhat higher than usual. This fact hints at the possibility that sequences in Greece decay more rapidly; but no conclusion can be reached with such a small sample size.

MAGNITUDE STABILITY IN TIME.

When the characteristics of aftershock sequences with respect to magnitude are examined, two sources of error are added to the possibility of the incompleteness of data; namely, lack of accuracy in magnitude determination, and confusion between different magnitude scales. Unfortunately, authors sometimes do not specify which scale they are using. When one single sequence is being examined, no problems arise, because the data are consistent within the sequence; if, however, results for different sequences are to be compared, the use of different magnitude scales may affect the conclusions. In the sequel, the various "local" magnitude scales, for the purposes of comparison of results among sequences, have been assimilated to M. Thus the only distinction left is between M and m; the latter scale has been used only in sequence (1).

Two aspects of the sequences have been examined in detail, namely, the variation of aftershock magnitudes with time, and the magnitude-frequency distribution. For all sequences except (1), where $1 \leqslant t \leqslant 100$ days, the data for $t \leqslant 100$ days have been included in the analysis. We first consider the distribution of magnitude with respect to time.

The overall mean magnitude, \overline{M} , has been calculated for each sequences as

$$\overline{M} = \frac{1}{K} \sum_{i=1}^{K} M_i$$
 , $i = 1, 2, ..., K$

where K is the total number of aftershocks in the sequence. Then the mean magnitude, M', of each group of 10 successive aftershocks is computed, thereby eliminating large individual fluctuations. In almost all the sequences considered the mean magnitude M' oscillates about M during the whole length of the sequence and no appreciable

decrease with time is detectable for $t \leq 100$ days. When a sequence lasts less than 100 days, the mean magnitude is stable throughout the sequence. Sometimes higher values of M' are observed in the first few hours after the main shock, but this can almost certainly be attributed to the fact that some shocks with $M \geq M^*$ are not detected when the frequency is very large. Figure 6 gives an example. The dotted line represents M.

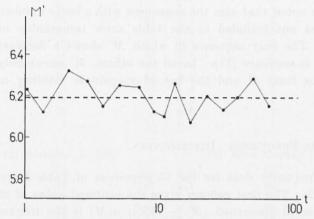


Fig. 6. - Magnitude stability in time: Aloutin Islands 1957.

Thus, the law of magnitude stability in aftershock sequences (first proposed by Lomnitz (14)) is fully confirmed; during an aftershock sequence the mean magnitude of the shocks is constant in time.

Sequence	М	Number of M'_i	% Within <i>M</i> ± 0.20
(1) Alaska	4.89	29	89.7%
(2) Aleutian I.	6.19	20	100 %
(5) Kern County	4.44	18	83.3%
(6) San Francisco	2.55	16	81.3%
(8) Parkfield	2.46	17	88.2%
(10) W. Thessaly	3.83	29	89.7%
(11) Amorgos	3.92	40	65.5%
(12) Magnesia	3.43	29	72.4%
(13) Zante	3.87	13	100 %
(14) Cremasta	3.75	10	100 %

Table VII - OSCILLATIONS OF MEAN MAGNITUDE.

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Consequently, the decrease of seismic activity with time is solely due to the decrease in aftershock frequency.

Moreover, the fluctuations of M' about the overall mean magnitude M are not very wide. Table VII summarizes the results for the sequences with a larger number of shocks. From left to right, the first column indicates the sequence, the second the overall mean magnitude, the third the number of calculated M'_i -points, and the fourth the percentage of such points which fall within the interval $\overline{M} \pm 0.20$. It is to be noted that also the sequences with a lesser number of aftershocks and not included in the table show remarkable magnitude stability. The only sequence in which M' shows a decreasing trend with time is sequence (11). In all the others, M' shows only random fluctuations from \overline{M} and the law of magnitude stability in time is satisfied.

MAGNITUDE-FREQUENCY DISTRIBUTION.

The frequency data for the 15 sequences of Table I are listed in Table VIII. The first column gives the centered value of magnitude in the interval concerned $(M \pm 0.05)$; n(M) is the frequency; and N(M) the cumulative frequency. For brevity, intervals in which the frequency was zero have been omitted from the Table. The magnitudes in sequence (1) are in the m-scale. The most commonly accepted form for the magnitude-frequency distribution, in case of both independent seismic events and aftershock sequences, is

$$\log n(M) = a - b M \tag{12}$$

where log is the logarithm to the base 10 and n(M) is the number of shocks with $M \pm dM$ (15).

Equation [12] is to be regarded as expressing a statistical relationship. Usually, the coefficients a and b have been calculated according to the least squares method. Suzuki (2) has pointed out that this is not rigorous, because $\log n(M)$ is not symmetrically distributed with uniform variance for each magnitude interval ΔM .

In this section we shall define n(M) is such a way that $\log n(M) = a$ not when M = 0, as in equation [12], but when $M = M^*$, where M^* is the minimum detectable magnitude in the sequence. Then the magnitude-frequency law takes the form

$$\log n(M) = a - b (M - M^*) , M \ge M^*.$$
 [13]

Table VIII - MAGNITUDE-FREQUENCY DISTRIBUTION.

M	n(M)	N(M)	М	n(M)	N(M)
	(1) Alaska 1964			(4) Desert Hot S _I	. 1948
4.5	44	294	3.0	15	68
4.6	45	250	3.1	14	53
4.7	32	205	3.2	9	39
4.8	36	173	3.3	4	30
4.9	30	137	3.4	6	26
5.0	26	107	3.5	5	20
5.1	24	81	3.6	3	15
5.2	9	57	3.7	$\overset{\circ}{2}$	12
5.3	13	48	3.8	Ī	iõ
5.4	12	35	3.9	$\frac{1}{2}$	9
5.5	10	23	4.3		7
5.6	7	13	4.4	3	6
5.7	3	6	4.5	i	3
5.8	$\frac{3}{2}$	3	4.6	BAR from day	2
5.9	Ĩ	ì	4.9	i	Ī
	(2) Aleutian I. 1957	1.8		(5) Kern County	1952
5.9	38	205	1.0		
6.0	49		4.0	25	184
6.1	29	167	4.1	29	159
$\frac{6.1}{6.2}$		118	4.2	26	130
	34	89	4.3	16	104
6.3	14	55	4.4	15	88
6.4	8	41	4.5	25	73
6.5	10	33	4.6	8	48
6.6	7	23	4.7	12	40
6.7	2 2 2 4	16	4.8	4	28
6.8	2	14	4.9	5	24
6.9	2	12	5.0	2	19
7.0		10	5.1	4	17
7.1	3	6	5.2	1	13
7.2	2	3	5.3	I	12
7.3	1	I	5.4	L	11
			5.5		10
	(A) T D 1 1000		5.6	Land to	9
	(3) Long Beach 1933		5.7	3	8
3.9	13	7.0	5.8	2	5
		78	0.1	$\overline{2}$	3
4.0	3	65	6.4	i i	1
4.1	$\frac{2}{17}$	62			
4.2	17	60			
4.4	16	43		(6) San Francisco	1957
4.5	3	27	20		
4.6	6	24	2.0	$\frac{25}{6}$	160
4.7	3	18	2.1	9	135
4.8	$\frac{2}{5}$	15	2.2	21	126
4.9	5	13	2.3	11	105
5.0	4	8	2.4	12	94
5.1	l	4	2.5	9	82
5.2	2	3	2.6	12	73
5.4			2.7	14	61

Table VIII - Continued

M	n(M)	N(M)	М	n(M)	N(M)
2.8	13	47	3.4	3	9
2.9	7	34	3.5	1	6
3.0	6	27	3.6	ļ	5
3.1	4	21	3.7	in the state of the state of	$\begin{smallmatrix}4\\3\\2\end{smallmatrix}$
3.2	4	17	4.1	$\frac{1}{1}$	3
3.3	3	13	4.5		1
3.4	3	10	5.0	le and	1
3.5	l	7			
3.6		6 5		(9) Chalkidike 1	932
3.8	2	9			0.5
4.0	1	$\frac{3}{2}$	3.4	5	85
4.2	1	1	3.5	$\frac{2}{2}$	80
4.4		1	3.6	2	78
			3.7	6	76
6	(7) Salinas 1963		3.8	13	70
1		10	3.9	12	57
1.0	2	46	4.0	5	45
1.1	$\frac{1}{4}$	44	4.1	10	40 30
1.3	4	40	4.2	TENT I THE TOTAL	23
1.4	l,	36	4.3	4	19
1.5	<u>l</u>	35	4.4	4 1	15
1.6	7	34	4.5		14
1.8	3	$\begin{array}{c} 27 \\ 24 \end{array}$	4.0	3	13
1.9	2	$\frac{24}{22}$	4.8	1	10
$ \begin{array}{c c} 2.0 \\ 2.1 \end{array} $	4	18	4.9	$\overset{1}{2}$	9
$\begin{bmatrix} 2.1 \\ 2.2 \end{bmatrix}$	3 2 4 2 2 4	16	5.0	$\overset{2}{2}$	7
2.2	4	14	5.1	ī	5
$ \begin{array}{c c} 2.3 \\ 2.4 \end{array} $	9	10	5.6		7 5 4 3
2.5	$\begin{array}{c}2\\2\\2\end{array}$	8	5.7	i	3
2.6	2	6	5.8	i	2
2.8		4	6.2	i i	ī
2.9		$\frac{4}{3}$	0.2		
3.1	i	$\overset{\circ}{2}$		(10) 11/ (11) 1	1054
4.6	i	ī		(10) W. Thessaly	1954
0.00	her is the line and a		3.2	13	299
	(8) Parkfield 1966		3.3	ì	286
158 498	(о) Рагкцета 1966	SE TV6	3.4	33	285
2.0	36	173	$\frac{3.4}{3.5}$	18	252
2.1	18	137	3.6	98	234
2.2	16	119	3.7	7	136
$\begin{bmatrix} 2.2 \\ 2.3 \end{bmatrix}$	20	103	3.8	8	129
2.4	13	83	3.9	34	121
2.5	15	70	4.0	15	87
2.6	9	55	4.1	9	72
2.7	7	46	4.2	17	63
2.8	5 7 10	39	4.3	7	46
2.9	7	34	4.4	11	39
3.0	10	27	4.5	6	28
3.1	5	17	4.6	6	22
3.2	2	12	4.7	2	$\begin{array}{c} 16 \\ 14 \end{array}$
3.3	1	10	4.8	2	14

Table VIII - Continued

М	n(M)	N(M)	М	n(M)	N(M)
4.9	4	12		(13) Zante 1962	
5.0	4	8	0.0		100
5.2	The state of the s	4	3.6	66	139
5.8	2	3	3.8	$\frac{12}{7}$	73
5.9	1	1	3.9	7	61
			4.0	29	54
	(11) Amorgos 1956		4.1	8	25
			4.2	$\frac{4}{2}$	17
3.5	87	400	4.3	2	13
3.6	78	313	4.4		11
3.8	22	235	4.5	3	10
3.9	78	213	4.6	1	7
4.0	40	135	4.7	1	7 6
4.1	12	95	4.8		5
4.2	27	83	4.9	j	5 4
4.3		56	5.2	r i	3
4.4	8	52	5.3	$\frac{1}{2}$	2
4.5	6	44	0.3		4
	6	38	77%		
4.6		32		(14) Cremasta 196	56
4.7	6 5		3.4	16	103
4.8		26	3.5	16 15	87
4.9	$\frac{4}{3}$	21	3.6	21	72
5.0		17	9.5	11	51
5.1	4	14	3.7	13	51
5.2	1	10	3.8		40
5.3	$\frac{2}{1}$	9	3.9	4	27
5.4		7	4.0	9	23
5.6	3	6	4.1	3	14
5.7	of me William to sai	3	$\frac{4.2}{4.3}$	3	11
5.8	1	2	4.3	1	8 7 4
6.8		1	4.4	3	7
13.89			4.6	$\frac{2}{1}$	4
	(12) Magnesia 1957		4.9	1	$\frac{2}{1}$
			5.1	1	1
3.0	53	291			
3.2	84	238		(15) Hawke's Bay 1	931
3.3	12	151			
3.4	73	139	4.1	2	71
3.5	7 4	66	4.2	10	69
3.6	4	59	4.3	3	59
3.7	10	55	4.4	6	56
3.8	11	45	4.5	10	50
3.9		34	4.6	2	40
4.0	5	31	4.7	5	38
4.1	3 5 2 2 1	26	4.8	2 5 2 4	33
4.2	$\frac{1}{2}$	24	4.9	4	31
4.2		22	5.0	5	27
4.4	3	21	5.1	5	22
4.5	4	18	5.2	3	17
4.6		14	5.3		14
1.7	$\frac{2}{1}$	12	5 4	$\frac{2}{3}$	12
4.7	I I	11	$\frac{5.4}{5.5}$	1	$\begin{array}{c} 12 \\ 9 \end{array}$
4.8	Enter By Lohan S	10	5.6		8
4.9 5.2	4				7
0.2		6	5.8	1	R
5.4	1	5	5.9	3	$\frac{6}{3}$
5.5	2	4	6.1		0
5.8		2	6.3		2
5.9	1		6.9		1
0.0			· · · ·		

Converting to natural logarithms one obtains

$$\ln n(M) = a' - b' (M - M^*)$$
 [14]

where

$$a' = \frac{a}{\log e} \quad , \quad b' = \frac{b}{\log e}$$
 [15]

Therefore

$$n(M) = e^{a'} e^{-b'(M-M^*)} = \gamma e^{-b'(M-M^*)}, e^{a'} = \gamma.$$

Normalizing, i.e., imposing the condition that

$$\int_{M^*}^{\infty} n(M) \, dM = 1$$

one has

$$\int_{M^{\bullet}}^{\infty} e^{-b'(M-M^{\bullet})} dM = \frac{\gamma}{b'} = 1$$

i.e., $\gamma = b'$. Thus we assume that the probability distribution of M takes the form

$$u(M) = b' e^{-h'(M-M^*)}, M \geqslant M^*.$$
 [16]

A procedure for estimating the parameter b' in [16] can be derived as follows. The mean of the distribution is

$$E(M) = \int_{M^*}^{\infty} M n(M) dM = M^* \div \frac{1}{b'}.$$

Approximating the population mean by the sample mean M, given by

$$\overline{M} = \frac{1}{K} \sum_{i=1}^{K} M_i$$

where K is the total number of shocks, one has

$$\overline{M} = M^* + \frac{1}{b'}$$

i.e.,

$$b' = \frac{1}{\overline{M} - M^*} \,. \tag{17}$$

The above procedure is but a particular case of the time-honored method of moments, introduced by K. Pearson and his school (24). Formula [17] was also proposed by Utsu at a meeting of the Seismological Society of Japan (reported by Aki (26)). The estimation of the parameter b' given by [17] has been proved by Aki (26) to be equivalent to the maximum likelihood estimate, and therefore has several desirable statistical properties.

Table 1X – Estimation of the parameter b in the magnitude-frequency law.

Sequence	N	N*	<i>b'</i>	ь	
(1) Alaska	4.89	4.45	2.273	0.987	
(2) Aleutian I.	6.19	5.85	2.941	1.277	
(3) Long Beach	4.40	3.85	1.818	0.790	
(4) Desert Hot Sp.	3.38	2.95	2.326	1.010	
(5) Kern County	4.44	3.95	2.041	0.886	
(6) San Francisco	2.55	1.95	1.667	0.724	
(7) Salinas	1.94	0.95	1.010	0.439	
(8) Parkfield	2.46	1.95	1.961	0.852	
(9) Chalkidike	4.13	3.35	1.282	0.557	
(10) W. Thessaly	3.83	3.15	1.471	0.639	
(11) Amorgos	3.92	3.45	2.128	0.924	
(12) Magnesia	3.43	2.95	2.083	0.905	
(13) Zante	3.87	3.55	3.125	1.357	
(14) Cremasta	3.75	3.35	2.500	1.086	
(15) Hawke's Bay	4.83	4.05	1.282	0.557	

Table IX gives the results of calculations. From left to right, the columns indicate sequence, sample mean \overline{M} , minimum magnitude M^* , and the estimations of the parameters b', b appearing in relations [16] and [13] respectively. The estimate of b' is given by [17], and b is given by [15]. The minimum magnitude M^* has been taken to be 0.05 units less than the values given in Table VIII because the value of M approximated to one decimal could actually come from

anywhere in the interval $M \pm 0.05$. It can be seen that b is usually slightly less than unity.

The basic idea for examining the observed fluctuations from the magnitude-frequency law is as in the case of the time distribution of aftershocks. Suzuki (2), among others, has argued that the number of shocks with $M \pm dM$ must follow a Poisson distribution. Accordingly, the problem is that of checking whether the observed deviation can be explained in terms of random fluctuations from the law.

It is more convenient to consider the cumulative distribution because individual large fluctuations in small intervals, possibly due to inaccurate magnitude determination, are smoothed out in this fashion, and moreover the normal approximation can be employed. The basic idea of the method is due to Suzuki (2). Some minor modifications have been introduced, and the normal approximation to the Poisson distribution, instead of the Poisson distribution itself, has been used.

From equation [16], it follows that the cumulative distribution function of magnitude has the form

$$F(M) = \int_{M}^{\infty} n(M) \ dM = e^{-b'(M-M^*)}.$$

Assuming that the total number of aftershocks in a sequence, K, coincides with the theoretical value for $M=M^*$, the expected value of the cumulative frequency at various $M>M^*$ is given by

$$N(M) = K e^{-b'(M-M^*)}$$
 [18]

which, when transformed by taking logarithms on both sides, is a straight line on semi-logarithmic paper, with slope equal to -b if the logarithms are to the base 10.

Now, to each magnitude range there corresponds a Poisson-distributed number of shocks. The cumulative number of shocks at a given magnitude M, therefore, is the summation of independent samples taken from each of the Poisson distributions corresponding to magnitudes greater or equal to M. According to the addition theorem for the Poisson distribution, such a cumulative number will also be Poisson-distributed. If the expected value is large enough, say, $N(M) \ge 10$ for all intervals, the Poisson distribution can be approximated by the normal distribution with mean N(M) and standard deviation $\sqrt{N(M)}$. It is then possible to calculate the fiducial in-

terval beyond which fluctuations are expected with a probability smaller than 5% (2.5% on each side). The limits of the interval such that

$$P\left\{\lambda_1 \leqslant N \leqslant \lambda_2\right\} \simeq 0.95,$$

are, for the normal distribution, $E(N) \pm 2 D(N)$; i.e., in the present case, $N(M) \pm 2 \sqrt{N(M)}$.

Accordingly, the above quantities have been calculated at

$$M_i = M^* + 0.1i$$
, $i = 1, 2, ..., k$.

Then, by joining all points of ordinates

$$N(M_i) + 2\sqrt{N(M_i)}$$
 and $N(M_i) = 2\sqrt{N(M_i)}$,

respectively, one obtains a confidence band which should contain approximately $95\,\%$ of the data points if they come from a population

Table X - VALIDITY OF THE MAGNITUDE-FREQUENCY LAW	Table	X	-	VALIDITY	\mathbf{OF}	THE	MAGNITUDE-FREQUENCY	LAW
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Sequence	Total number of data points	Number of points outside conf. limits	% Points outside
(1) Alaska	15	52000 4 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	26.7%
(2) Aleutian I.	11	0	_
(3) Long Beach	12	2	16.7%
(4) Desert Hot Sp.	9	0	<u> </u>
(5) Kern County	15	0	_
(6) San Francisco	17	0	<u> </u>
(7) Salinas	16	0	_
(8) Parkfield	15	0	-
(9) Chalkidike	17	2	11.8%
(10) W. Thessaly	24	14	58.3%
(11) Amorgos	18	1	5.6%
(12) Magnesia	17	3	17.6%
(13) Zante	9	2	29.2%
(14) Cremasta	10	0	
(15) Hawke's Bay	16	0	<u>-</u>

whose expected value is given by the magnitude-frequency law as expressed by [18]. The rightmost interval (say, from M_I to ∞) has always be chosen in such a way that $N(M) \ge 10$.

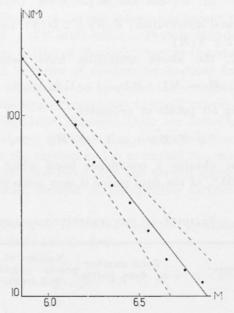


Fig. 7. - Magnitude-frequency distribution: Aleutian Islands 1957.

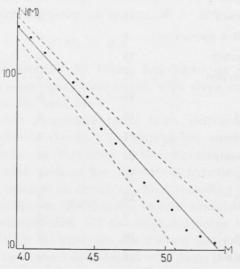


Fig. 8. - Magnitude-frequency distribution: Kern County 1952.

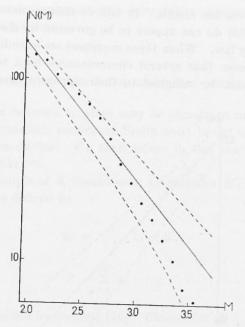


Fig. 9. - Magnitude-frequency distribution: San Francisco 1957.

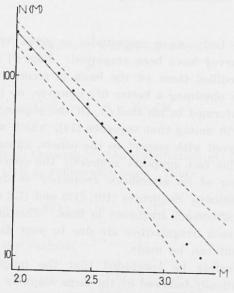


Fig. 10. - Magnitude-frequency distribution: Parkfield 1966.

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Table X gives the results. It follows that sequences (1), (3), (9), (10), (12) and (13) do not appear to be governed by the assumed magnitude-frequency law. When these sequences are examined one by one, however, it is seen that several circumstances tend to decrease the weight that must be assigned to their apparent irregularity. For

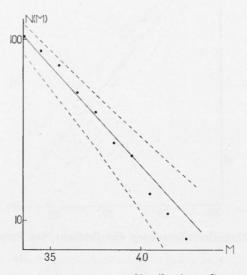


Fig. 11. - Magnitude-frequency distribution: Cremasta 1966.

sequence (1) the body—wave magnitudes as given by the U.S. Coast and Geodetic Survey have been employed. Page (3) found them inaccurate and modified them on the basis of records of five selected stations, thereby obtaining a better fit; however, no test on the fluctuations was performed in his study. Of the sequences occurring in Greece, it is worth noting that sequence (14), which was studied with a network improved with respect to the others, appears to follow the assumed law; this fact supports indirectly the conclusion that the irregular behavior of its aftershock frequency is due to changing local stress conditions. Sequences (10), (12) and (13) are also irregular with respect to aftershock frequency in time. This fact points to the likelihood that such irregularities are due to poor data, although no definite statement can be made.

It can therefore be concluded that the magnitude-frequency law [18] is statistically followed by the large majority of the sequences under consideration. Figures 7 to 11 give some examples; the logarithmic ordinate represent the cumulative frequency, the linear abscissa the magnitude.

CONFIDENCE LIMITS ON THE FREQUENCY PARAMETER.

Differences between b-values may be significant or not. In order to decide the question, confidence limits must be set on the estimated value of the parameter. We shall follow in this matter a procedure suggested by Λ ki (26).

Given a sample of K shocks with magnitudes M_i , $i = 1, \ldots, K$, let y_i and Y be defined by

$$y_i = \frac{\delta}{\delta b'} \ln n(M_i)$$

$$Y = \sum_{i=1}^K y_i$$

where n(M) is given by formula [16]. Clearly for all i,

$$y = \frac{1}{b'} - M + M^*$$

and therefore

$$E(y) = \int_{M^*}^{\infty} y \ n(M) \ dM = 0$$

$$V(y) = \int_{M^*}^{\infty} y^2 \ n(M) \ dM = \frac{1}{b'^2}$$

By the central limit theorem $(^{24})$ the distribution of Y will be approximately normal if K is sufficiently large. Since

$$E(Y) = 0 \quad , \quad V(Y) = \frac{K}{{h'}^2}$$

it follows that the variable

$$Y^* = \frac{Y - E(Y)}{D(Y)} = \frac{b' Y}{\sqrt{K}}$$

is approximately normally distributed with mean 0 and standard deviation 1; therefore

$$P \, \left\{ \, -1.96 \leqslant Y^* \leqslant 1.96 \, \right\} \, = \, 0.95 \; .$$

Accordingly, the 95% confidence limits on b' are obtained by solving the inequality

$$-1.96 \leqslant \frac{b' \ Y}{\sqrt{K}} \leqslant 1.96$$

which gives

$$-1.96 \leqslant \frac{b'}{\sqrt{K}} \sum_{i=1}^{K} \left(\frac{1}{b'} - M_i + M^* \right) \leqslant 1.96$$

i.e.,

$$rac{1-1.96 \, / \, \sqrt{K}}{M-M^*} \, \, \leqslant \, \, b' \, \, \leqslant \, \, rac{1\, + \, 1.96 \, / \, \sqrt{K}}{M-M^*}$$

Table XI - Confidence limits on the frequency parameter.

Sequence	95% Confidence limits		
(2) Aleutian I.	$1.102 \leqslant b \leqslant 1.452$		
(4) Desert Hot Sp.	$0.770 \leqslant b \leqslant 1.250$		
(5) Kern County	$0.757 \leqslant b \leqslant 1.015$		
(6) San Francisco	$0.612 \leqslant b \leqslant 0.836$		
(7) Salinas	$0.312 \leqslant b \leqslant 0.566$		
(8) Parkfield	$0.725 \leqslant b \leqslant 0.979$		
(11) Amorgos	$0.833 \leqslant b \leqslant 1.015$		
(14) Cremasta	$0.876 \leqslant b \leqslant 1.296$		
(15) Hawke's Bay	$0.427 \leq b \leq 0.687$		

Table XI summarizes the results for the 9 sequences where the assumed magnitude-frequency law appears to hold. According to usage, the parameter b, instead of b', has been employed. It can be seen that the b-values cluster around the interval 0.8-1.0.

Although sequence (2), which occurred along an active island are, shows an anomalously high value of the frequency parameter, the sample size is too small to support the contention that b has some geotectonic significance.

CONCLUSION.

The basic statistical laws describing the phenomenology of aftershock sequences are confirmed by the present study. These laws are as follows:

- (1) Aftershock sequences are generally crustal events, although deeper ones also occur;
- (2) The frequency of aftershock occurrence within the same sequence decays in time according to the law

$$n(t) = \alpha t^{-\beta} \quad , \quad t \geqslant 100$$

where the decay parameter β is approximately equal to, or slightly greater than, unity;

- (3) The aftershock magnitudes, apart from individual fluctuations, show stability in time to the end of the sequence;
- (4) The frequency-distribution of magnitude in a sequence is of exponential form

$$N(M) = K e^{-b'(M-M^*)}$$

where the frequency parameter $b=b'\log e$ is usually slightly less than unity

The importance of the mode of grouping the data in a statistical analysis makes it desirable to introduce a standardized procedure. Furthermore, an appropriate statistical method must be employed, and the hypotheses involved must be checked. A systematic treatment of data greatly increases the reliability of the results.

It also turns out that the least squares method, when the observed values are uncorrelated, yields rather satisfactory results even if the conditions for its theoretical validity are not met.

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REFERENCES

- (1) PAGE R., Focal Depths of Aftershocks, "J. Geophys. Res.", 73, 3897-3903 (1968).
- (2) Suzuki Z., A Statistical Study of the Occurrence of Small Earthquakes, III, "Sci. Rep. Tohoku Univ., 5th Series, Geophys.", 10, 15-27 (1958)
- (3) PAGE R., Aftershocks and Microaftershocks of the Great Alaska Earthquake of 1964. "Bull. Seism. Soc. Am.", 58, 1131-1168 (1968).
- (4) DUDA S. J., Phänomenologische Untersuchung einer Nachbebenserie aus dem Gebiet der Aleuteninseln. "Freib. Forschungshefte". C-132, 5-90 (1962).
- (5) BENIOFF H., Earthquakes and Rock Creep, I. "Bull. Seism. Soc. Am.", 41, 31-62 (1951).
- (8) RICHTER C. F., ALLEN C. R., and Nordquist J. M., The Desert Hot Springs Earthquakes and Their Tectonic Environment, "Bull. Seism. Soc. Am.", 48, 315-337 (1958).
- (7) RICHTER C. F., Foreshocks and Aftershocks, in Earthquakes in Kern County, California, during 1952, "Bull. Cal. Div. Mines.", 171, 177-197 (1955).
- (8) TOCHER D., Seismographic Results from the 1957 San Francisco Earthquake, in San Francisco Earthquakes of March 1957, "Cal. Div. Mines Sp. Rep.", 57, 59-71 (1959).
- (9) Udias A., A Study of the Aftershocks and Focal Mechanism of the Salinas-Watsonville Earthquakes of August 31 and September 14, 1963, "Bull. Seism. Soc. Am.", 55, 85-106 (1965).
- (10) McEvilly T. V., Bakun W. H., and Casaday K. B., The Parkfield, California, Earthquakes of 1966, "Bull. Seism. Soc. Am.", 57, 1221-1244 (1967).
- (11) PAPAZACHOS B., DELIBASIS N., LIAPIS N., MOUMOULIDIS G., and PURCARU G., Aftershock Sequences of Some Large Earthquakes in the Region of Greece, "Ann. Geofisica", 20, 1-93 (1967).

- (12) COMNINAKIS P., DRAKOPOULOS J., MOUMOULIDIS G., and PAPAZACHOS B., Foreshock and Aftershock Sequences of the Cremasta Earthquake and Their Relation to the Waterloading of the Cremasta Artificial Lake, "Ann. Geofisica". 21, 39-71 (1968).
- (13) Mogi K., On the Time Distribution of Aftershocks Accompanying the Recent Major Earthquakes in and near Japan, "Bull. Earthq. Res. Inst.". 40, 107-124 (1962).
- (14) LOMNITZ C., Magnitude Stability in Earthquake Sequences, "Bull. Seism. Soc. Am. ", 56, 247-249 (1966).
- (15) GUTENBERG B., and RICHTER C. F., Seismicity of the Earth and Associated Phenomena, 2nd ed., Princeton Univ. Press., Princeton (1954).
- (18) MATUZAWA T., Study of Earthquakes, Uno Shoten, Tokio (1964).
- (17) BATH M., and BENIOFF H., The Aftershock Sequence of the Kamchatka Earthquake of November 4, 1952, "Bull. Seism. Soc. Am.", 48, 1-15 (1958).
- (15) Iosif T., and Radu C., Elastic Strain Characteristics in the Deep-Focus Earthquakes of the Vrancea Region, "Stud. Cerc. Astr. Seism.", 2, 269-285 (1961), [in Romanian].
- (19) LUKK A. A., The Aftershock Sequence of the Dzhurm Deep-Focus Earth-quake of March 14, 1965, "Izv. Earth. Phys.", 5, 83-85 (1968) [in Russian].
- (20) JEFFREYS Sir H., Aftershocks and Periodicity in Earthquakes, "Gerl. Beitr. Geophys.", 53, 111-139 (1938).
- (21) UTSU T., On the Nature of Three Alaskan Aftershock Sequences of 1957 and 1958, "Bull. Seism. Soc. Am.", 52, 279-297 (1962).
- (22) GRAYBILL F. A., An Introduction to Linear Statistical Models, Vol. I, McGraw-Hill, New York (1961).
- (23) YULE G. U., and KENDALL M. G., An Introduction to the Theory of Statistics, 12th ed., Griffin, London (1940).
- (24) Cramér II., Mathematical Methods of Statistics, Princeton Univ. Press, Princeton (1946).
- (25) RALSTON A., A First Course in Numerical Analysis, McGraw-Hill, New York (1965).
- (26) AKI K., Maximum Likelihood Estimate of b in the Formula log N = a b M and its Confidence Limits, "Bull. Earthq. Res. Inst.", 43, 237-239 (1965).